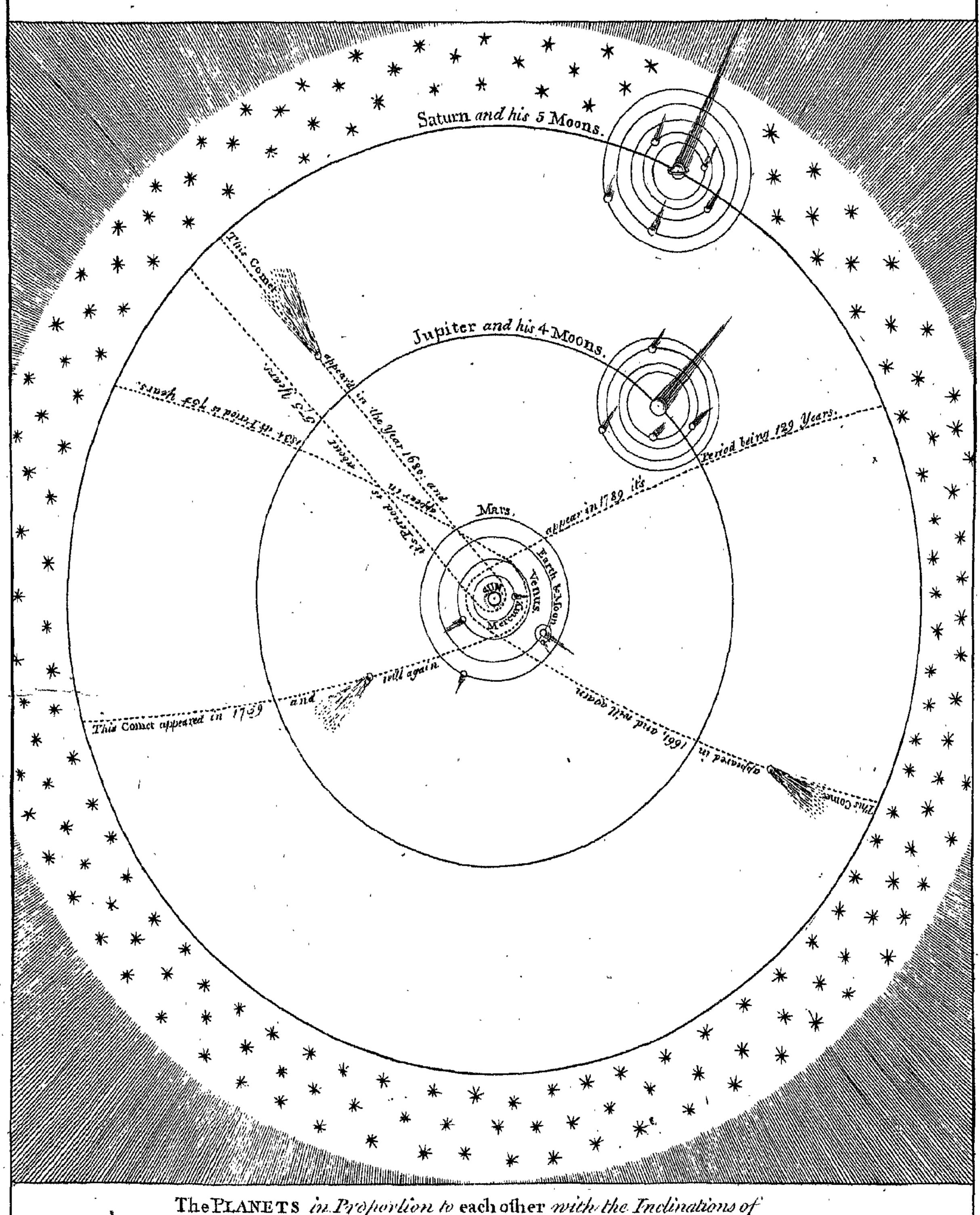
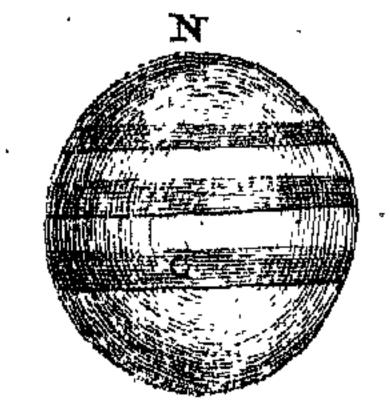
The COPERNICAN or TRUE SYSTEM of the UNIVERSE:

exhibiting the Orbits of the PLANETS according to their mean Distances from the SUN.



The PLANETS in Proportion to each other with the Inclinations of their Axes &c.



JUPITER and his Belts.

MARS.

EARTH.

VENUS.

MERCURY.

MOON.

ATURN and his Ring.

Note: The Sun is about 10 inches in Diameter, according to the above Proportion of the Flancts.

VIEW of the HEAVENS:

Being a Short, but Comprehensive,

SYSTEM

O F

MODERN ASTRONOMY.

EXHIBITING,

- I. The Number, Order, Distances, Magnitudes, and Periods of all the Planets and their several Moons, composing our System, which the Learning of the present Age esteems as so many Worlds full of Inhabitants.
- II. The length of the Day and Year, with the Variety of the Seasons in each Planet; and alfo the Phænomena of the Heavens to the Inhabitants thereof.
- III. Some Account of the Comets, their Number, Period, and Appearances; and also the Directions of their fiery Trains thro' the Heavens; with probable Conjectures of the Uses of those amazing Bodies.
- IV. The Number, Magnitude, and Distances of the Fixed Stars; with their Divisions into Signs, Catalogues, and Constellations.
- V. The true Figure of the Planets' Orbits, with the Nature of the Motions in them; their

- Aphelions, Perihelions, Eccentricities, Nodes, &c. &c.
- VI. The Inequality in the Length of the Natural Day, commonly called the Equation of Time.
- VII. The Power that retains the Earth and Planets in their Orbits.
- VIII. The direct and retrograde Motions of all the Planets; as also, the Reasons why they sometimes appear stationary, or not to move at all.
- IX. The Nature and Causes of Eclipses, both of the Sun and Moon; with an easy and expeditious Method of calculating the Eclipses which will happen in any Year.
- X. The Description and Use of a curious Astronomical Clock, which will shew the Hour of the Night by the Stars.

TO WHICH IS ADDED,

The Use of the CÆLESTIAL GLOBE:

WITH

Its Application to a Number of very interesting Problems. Concluding with some curious Phænomena upon the Sun and Moon exhibited in a darkened Room; and a sew select Paradoxes, intended to excite the Attention of the Learner.

The whole illustrated with Copper-plates of the System, the Sun, Moon, Eclipses, &c. and difposed in so easy and natural a Manner, as to be understood in a few Days.

THE SECOND EDITION, WITH MANY ADDITIONS AND IMPROVEMENTS.

By the Rev. Mr. TURNER, of Magdalen-Hall, Oxford;

Rector of Comberton; —Vicar of Elmley; —Minister of Norton; —and Chaplain to the Right Honourable the Countess Dowager of Wigton. —Author of The Heavens Survey'd; —The View of the Earth; —Plain Trigonometry rendered Easy and Familiar; —System of Gauging; —Chronologer Perpetual; —and a New Introduction to Book-keeping.

The WORLDS were framed by the Word of God.

PAUL.

LONDON:

Printed for S. CROWDER, at No. 12, Pater-noster-Row. MDCCLXXXIII.

THE

HEAVENS SURVEY'D,

ANDTHE

True System of the Universe

Delineated, so as to form a Curious

ASTRONOMICAL INSTRUMENT:

EXHIBITIN.G

The Number, Order, Periods, Aphelions, Perihelions, Nodes, &c. of all the PRIMARY PLANETS by Inspection:

TOGETHERWITH

An easy and expeditious Method of ascertaining by the same Instrument, and a new Set of Tables, their Anomalies, Longitudes, Latitudes, Retrogradations, Conjunctions, Elongations, and Distances at all Times, both Heliocentric and Geocentric; that is, as seen from the Sun and from the Earth; and also from one another.

In the Course of the Work is met with

A View of the System from the Earth, elucidating and explaining the true Causes of the Directions and Retrogradations of the Planets; and of the various Phases they exhibit to us here: With the Reasons of the Transits of Venus and Mercury over the Sun; the Times when they happen;—and an easy Method of delineating the Tracks those Planets take in passing over his Dish.——Also, the remarkable Path which the Planet Mars seemed to describe in the Heavens in the year 1762.

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And constructed in a Manner so easy and natural, as to convey to the Astronomical Learner a perfect Knowledge of the Solar System at first View.

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Author of the View of the Heavens; -----View of the Earth; ----- Trigonometry rendered Easy and Familiar; ----- System of Gauging; ----- Chronologer Perpetual; ----- and a New Introduction to Book-Keeping.

Rector of Comberton; Vicar of Elmley; Minister of Norton, and Chaplain to the Right Honourable the Countess Dowager of Wigton.

Cælique Meatres

Describent Radio; et surgentia Sidera dicent.

VIRG.

LONDON:

Printed for S. CROWDER, at No. 12, Pater-noster-Row: MDCCLXXXIII.

THE RIGHT HONOURABLE

FREDERICK, LORD NORTH, L.L.D.

CHANCELLOR;

TO THE REVEREND

The VICE-CHANCELLOR;

AND TO THE

HEADS OF HOUSES,

AND

MEMBERS OF CONVOCATION,

INTHE

UNIVERSITY of OXFORD,

THIS

COMPENDIUM OF ASTRONOMY,

(Drawn up with a View to render the first Rudiments of so illustrious and useful a Science more easy and practicable to the Minds of Youth, and particularly of those who receive their Education in that ancient and venerable Seat of Learning)

IS

WITH THE MOST PROFOUND HUMILITY ADDRESSED,

A N D

THEIR PROTECTION AND FAVOUR MOST RESPECTFULLY CRAVED,

By THE AUTHOR.

A SHORT

SYSTEM

O F

MODERN ASTRONOMY.

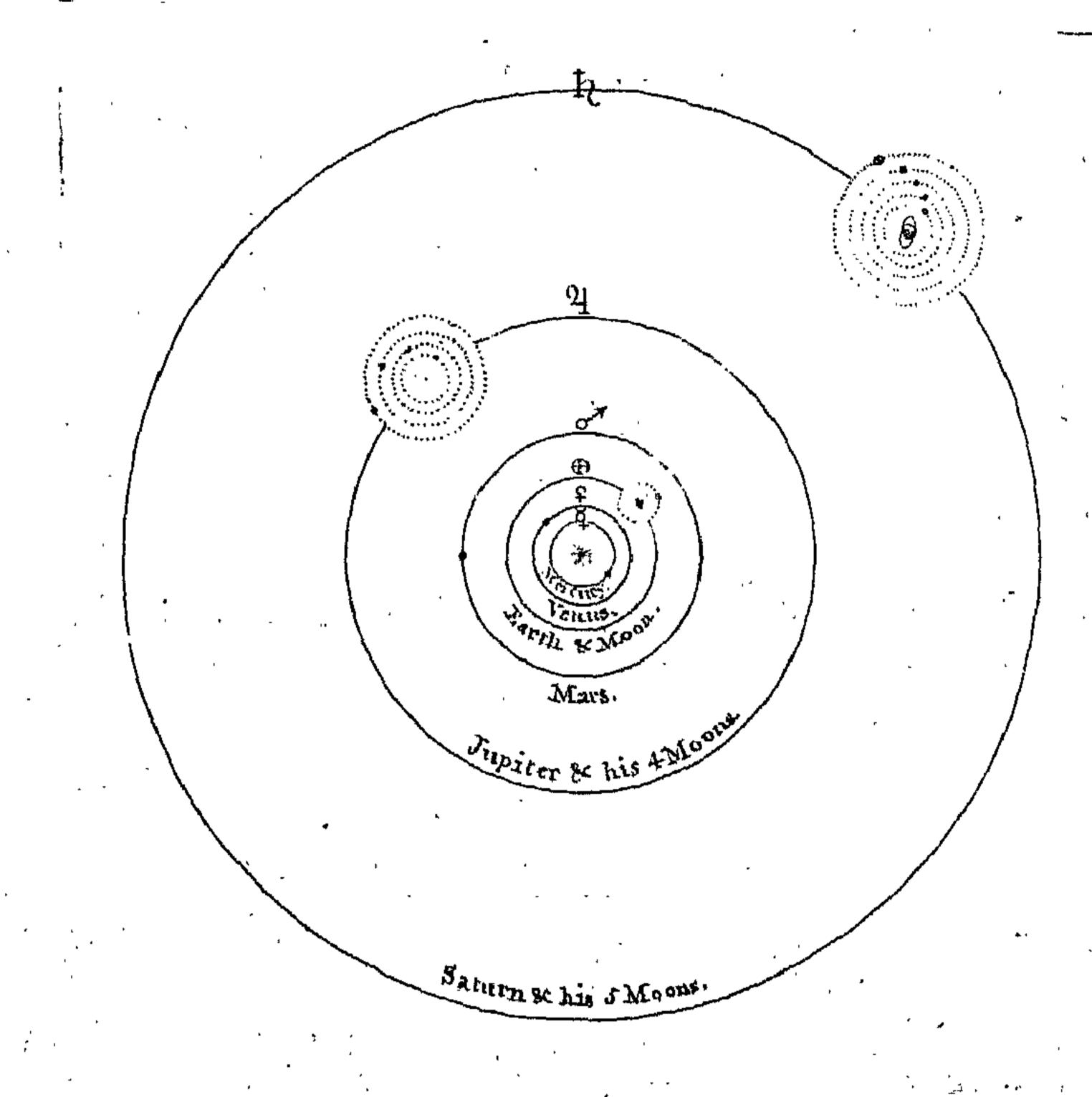
STRONOMY is that Part of Learning, which contemplates the Number, Order, Distances, Magnitudes, Periods, and Appearances of the Heavenly Bodies; as the Sun, Moon, Planets, Comets, and Stars,—And is a Knowledge, which contributes not only to the Enlargement of our Ideas of the Immensity, Magnisicence, and transcendent Grandeur of God; but also affords the sublimest and most noble Entertainment to the Understanding and Mind of Man.

When we look up to the Heavens, which Way soever we turn our Eyes, some of those wonderful Luminaries present themselves to our View: And, as God hath old us—that He ordained them for Signs and for Seasons; for Days and for Years;—it is undoubtedly our Duty to observe their appointed Periods, that they may answer the great end—so wise—so beneficent a Providence intended them for.

As TRONOMERS, in Consequence of this divine Appointment, to determine the Times and Seasons,—have found, by their repeated Observations, that several bodies move round the Sun: And, as they appear through the Telescope to be large Globes of Earth, like our own, (and, that our Earth, viewed from them, would appear nearly as they do to us) have justly concluded them to be Worlds enlighten'd and warm'd by the Sun, as ours is, and inhabited by various Species of Beings; though perhaps very different from those on our Globe; but properly form'd and constituted, for the Situation God hath placed them in there.

The

The Names of the several Bodies, (which we call Planets) found moving round the Sun, and composing our System, together in the Order in which they revolve, are—First, Mercury ($\stackrel{\triangleright}{}$)*—next, Venus ($\stackrel{\triangleright}{}$)—the Third, our Earth ($\stackrel{\bigcirc}{}$) with the Moon ($\stackrel{\triangleright}{}$)—the Fourth, Mars ($\stackrel{\triangleright}{}$)—the Fifth, Jupiter ($\stackrel{\triangleright}{}$) and his 4 Moons—the Sixth, Saturn ($\stackrel{\triangleright}{}$) and his 5 Moons: All which are represented at their proper Distances, with respect to each other, from the Sun, in the Scheme at the Beginning of the Book; or in This here delineated.



What Kind of System was received in the first Ages of the World, has not been transmitted to us; but this System we find taught and profess'd by Pythagoras, a learned Greek Philosopher, about 500 Years before Christ. After his Time, it lay dormant, till Nicholas Copernicus, a Clergyman in Prussia, reviv'd it about the latter End of the 15th Century; and now it has gain'd the Esteem of the Learned World +.

In this System we see, as in a Glass, the Beauties and Harmony of the Universe display'd; for here are neither solid Orbs nor Chrystaline Heavens to carry the Planets round, as vainly imagined by Ptolemy and Others; but every Thing appears in a most simple, rational, and demonstrable Order.

It is made up and adorn'd with 17 Bodies, which we shall immediately proceed to speak of, and in the Order and Situation they are found in the System itself.

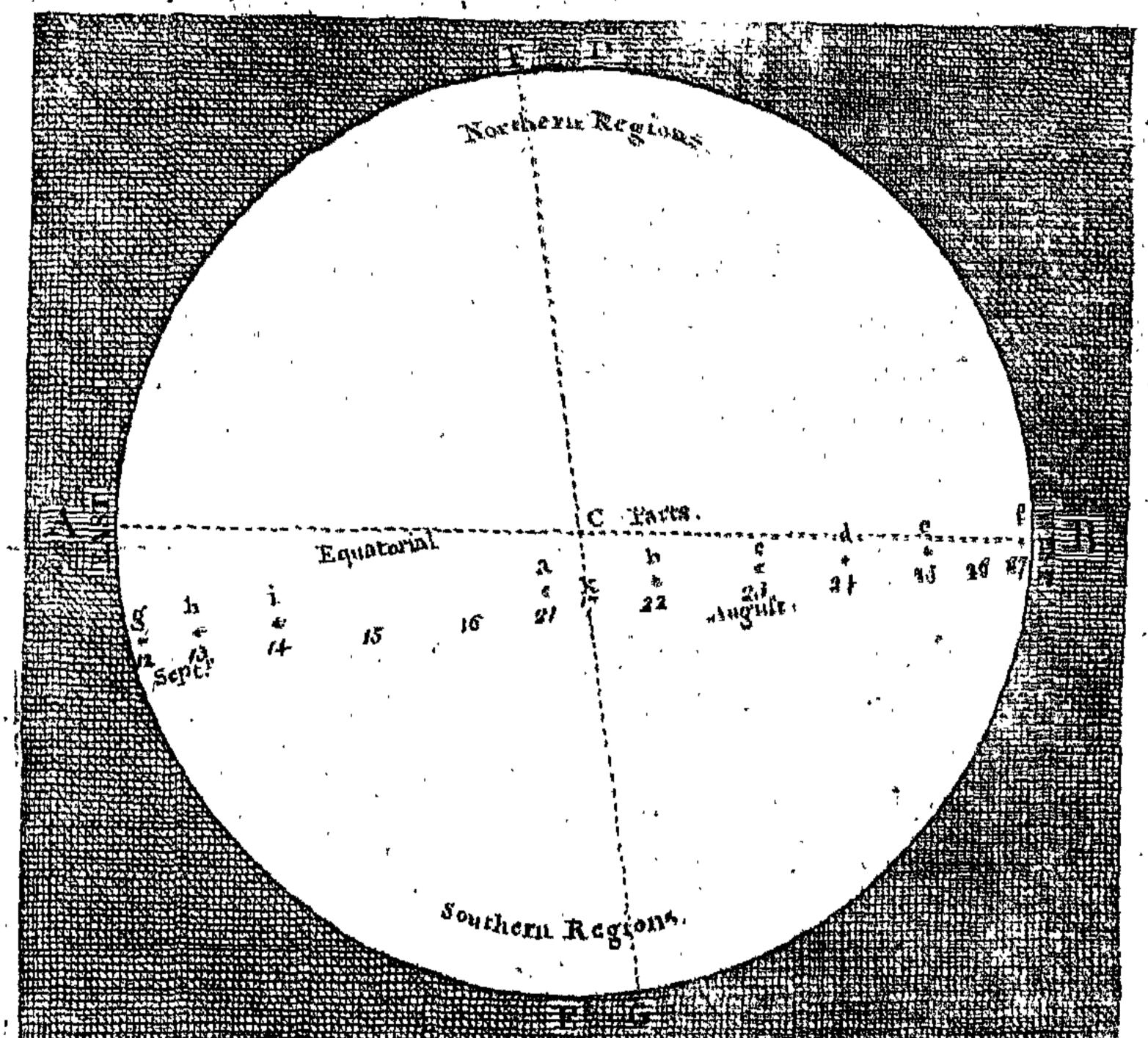
*The Characters placed with the Names of the Planets, are, for Brevity's sake, commonly used by Astronomers, instead of the Words at Length; as & for Mercury; 2 for Kenus, &c.

+ Most of the Ancients (not acquainted with Mathematical Learning) took it for granted that the Earth flood fill, and the Sun moved, because it appeared to do so. Accordingly, their most samous System, published by Ptolemy, an Egyptian Astronomer, about 138 Years after Christ, supposes the Earth immoveably fixed in the Center of the Universe, and the 7 Planets, viz. Moon, Mercury, Venus, the Sun, Mars, Jupiter, and Saturn, to revolve in Circles round it. Beyond these is placed the Firmament of the fixed Stars; then two Chrystaline Spheres: All which were included in, and received Motion from the Primum Mobile, which constantly carried all those vast and solid Orbs about the Earth in 24. Hours, from East to West.—But this System being found to be too much accommodated to Sense to stand the Test of Art; others were contrived and published at different Times; but all of them being embarrassed with many Difficulties and Absurdities, were embraced by a few; and at last were obliged to give Wuy to the only true and rational Solar System, restored by Copernicus, as mentioned above.

Of the Sun.

IRST, we find the Sun in the Center of the System, immoveable from thence, having no circular Motion in an Orbit, but a central one only about his own Axis, in the Space of 25 Days and a Quarter, as discover'd by viewing him through a Telescope. He is about 800.000 Miles in Diameter; is one Million of Times larger than our Earth; and by his rotation dispenses Light, Heat, and Motion, to all the Planets revolving round him.

By observing the Sun through a good Glass, he generally appears to have one or more black Spots on his Body, which come on first on the eastern Side, pass over his Face, and go off on the western, in a little more than 13 Days, and in the same Space of Time return again. By these Spots not only his Rotation on his Axis was discovered as above, but also that his Axis leans or inclines to the Orbit of the Earth in an Angle of about 82 Degrees. The Places, and Manner in which a large Spot appear'd to pass over the Sun's Disk, in August and September, 1764, are as here delineated.



On the 21st of August the Spot was seen at a; on the 22d at b; the 23d at c; 24th at d; on the 25th it was at e; the 26th not observed; on the 27th it appeared at f; on the 28th not observed; on the 29th supposed to be on the western Edge. After 13 Days Absence, i. e. on Sept. 12th, the Spot appeared again on the eastern Edge at g; on the 13th Day it was at b; on the 14th at i; the two following Days were cloudy; but on the 17th it was found at k; having completed a Revolution.—F G represents the Axis of the Motion of the Spot. F the North-pole, and G the South-pole. ACB is the Ecliptic, or Tract opposite the Earth's Orbit. D and E its two Poles. This Spot was about a thirtieth Part of the Diameter of the Sun, consequently was near 26.000 Miles in Length; its Breadth was something less.—A Spot, less in Diameter than our Earth, is not visible without a good Glass*.

* These Solar Spots do not always remain the same, but sometimes old ones vanish, and afterwards others succeed in their Room; sometimes several small ones gather together and make one large Spot; and sometimes a large Spot is seen to be divided into many small ones. But, notwithstanding these Changes, they all turn round with the Sun (towards the West), in the same time, which is an undeniable Proof they are on his Surface and not at a Distance from him. Whilst

Whilst the Sun turns thus on its Axis in 25 Days and a Quarter, it throws off from itself a fine subtle Matter, which is extended through all Parts of the System. This Matter is what constitutes Light; and comes with such Rapidity, that it arrives from the Sun hither in seven or eight Minutes Time. It passes through the Pores of Glass and other transparent Bodies, with little or no Resistance; but when it falls upon Bodies it cannot penetrate, it is restected to the Eye, and so renders them visible. By this Means, we know that the Planets are opaque, or dark Bodies in themselves, and shine only by this reslected Light; for when the Sun is in such a Position that it cannot shine upon that Part of them next us, they become obscure and invisible.

There are fix Primary or principal Planets, und ten Moons, which receive their Light from the Sun, viz. Mercury, Venus, the Earth and Mooon, Mars, Jupiter and his four Moons, Saturn and his five Moons. Each Primary Planet describes a large Orbit round the Sun, and being placed at different Distances, one beyond the other, make their Revolutions in different Periods of Time.

In these Revolutions of the Planets, there is one thing very remarkable, which is, that they are all made the same Way, i. e. from West towards the East, and opposite the middle or equatorial Parts of the Sun; which caused some Philosophers to imagine they were carried about by a Vortex or Whirlpool of Æthereal Matter. But many Observations made on the Heavenly Bodies by the Moderns, have sufficiently confuted that Opinion *.

^{***} If the young Astronomer has a Mind to observe the Spots upon the Sun; the best Way will be to use a Refracting Telescope of about six or eight Feetor a Reflecting one of two or three, with a smeaked Glass placed before the Eyeglass next the Eye, which will take off the glaring Light, and render the Spots visible. By this Method their Appearances may be observed, Day by Day, and the Tract they describe in their Passage over the Sun's Disk easily ascertained.—Or, the Image or Picture of the Sun, with its Spots, may be received into a dark Room, through a Telescope of one or two Feet long, (without a smoaked Glass) upon a Piece of white Paper, which may be magnified or diminished, by bringing the Paper nearer or further from the Glass, to the Dimensions you please. This is an innocent and easy Way; but one Inconvenience attends it, which is, that the Image, seen upon the Paper, always appears in an inverted Position.

^{*} This will be further considered, when we come to treat of the Motion of Comets.

MODERN ASTRONOMY.

Of the Planet Mercury.

Degrees: And, because the Heavens, at the Time Mercury is at his greatest apparent Distance from the Sun, are so illuminated, there can be no Observation made to discovered.—The Diameter * of this Planet is about 2460 Miles; which makes him near 30 Times less in Bulk than our Earth.

The Year †, to the Inhabitants of this Planet; (for it is the Opinion of the Learned, that the several Planets are so many Worlds, surnish'd with Beings of different Kinds as our Earth is,) is not quite so much as one Quarter of Ours: But as they are almost three Times nearer the Sun than we are, his Face must consequently appear nearly three Times bigger; and his Light and Heat nearly nine Times greater than with Us. This Degree of Heat would continually keep our Waters boiling, and render it impossible for Us to live; but the Bodies of Animals and Vegetables there, are, no doubt, properly tempered to sustain it. They, as well as their Planet, may be constructed more dense, firm, and compact than we are here; and may require that great Degree of Heat to support them in Life, which would destroy Beings of our softer Texture, and consume them away.

The Length of the Day, and the Variety of the Seasons there, is totally unknown to our Astronomers at present: For they have not been able (by Reason of his Nearness to the Sun, and the great Illumination of the Heavens, when at his greatest Distance from him) to discover the Inclination of his Axis, or the Time he revolves about it.

The People in this Planet will observe the Spots on the Sun's Disk much plainer than we can, and, by that means, be enabled to discover his Rotation on his Axis, and better ascertain what those Spots are. They will see all the other Planets revolve round them: Consequently Venus and our Earth, when opposite the Sun, will shine with a full Face towards them, and afford a great Light by Night there; but Mars, Jupiter, and Saturn will appear nearly the same as they do to us, only they will not give them quite so much Light on Account of their greater Distance from that Planet.

The Distances of the Planets from Us, and from one another; and also their Diameters, are obtained by Trigonometry; -- An easy Method of making those Calculations, is shewn in my Treatise on that Subject. The Distances of the Planets from the Sun given in this System are what were computed by the Astronomers till the late Transit of Venus; but from very accurate Observations made on that Phanomenon, their Distances are now found to be something above I Part in 6 more. Both Distances are expressed in the Table at Page 27.

† The Revolution of a Planet round the Sun, is called the Annual or Yearly Motion, because the Inhabitants have all the Alterations of the Seasons, and Varieties of the Year complete in that Revolution.

---And the Rotation of a Planet round its Axis, is called the Diurnal or Daily Motion; because by this Means, each part of its Surface is carried successively towards and from the Sun, which always illuminates that Half next him, making it Day; whilst the other Half remaining in Darkness, must

The Periods of the Planets are determined by observing the Time of their Departure from a fix'd Star, till they arrive to the same Star again.—And the Diurnal Rotation round their Axis is discovered by Spots seen with the Telescope upon their Disks, which Spots appear to pass over their Faces in a certain Space of Time, and after being absent the same Time, return again and pass over their Disks as before.

Of

Of the Planet VENUS.

TEXT beyond Mercury is the glittering Planet Venus: She makes her periodical Revolution in 224 Days, 16 Hours, and 49 Minutes, at the Diftance of 59 Millions of Miles from the Sun. She is observed to turn upon her Axis in 23 Hours, and her Diameter is about 7906 Miles, which is nearly equal - to that of our Earth.

This Planet, as well as Mercury, can never be seen at Midnight; but is visible only three or four Hours in the Morning or Evening, according as she is before or after the Sun. When she is in that Part of her Orbit West of the Sun, she rises in the Morning before him, and is called the Morning-star. When she is on the east Side of the Sun, she sets in the Evening after him, and is then the Evening star.

Since these two Planets, Venus and Mercury, always accompany the Sun, (and are never farther from him than, the latter about 27 Degrees, and the former 47,) Astronomers well know, that the Earth's Orb must circumscribe their Orbs; and that they revolve about the Sun in Circles much nearer than we are: Hence it is, they are never found one fixth or scarcely one-eighth Part of a whole Circle from him; which they would be, if the Earth was in the Center of the System, and they went round us *.

To the Inhabitants of this Planet, the Sun will appear almost twice as big as he does to us: His Face, and consequently his Light and Heat, must be almost four Times greater. The Year in that Planet is about two-thirds of our Year, and the Length of the Day 23 Hours.

They who observe the Heavens there, will see four Planets above them; viz: the Earth, Mars, Jupiter, and Saturn; and one below, which is Mercury. When our Earth is in Opposition to the Sun, it will appear to shine (in the Night) there with a full Face, and very bright. The Moon will always feem to wait upon the Earth, and never appear more than half a Degree (or two Hands Breadth) from it. Mercury will accompany the Sun, and be seen as a Morning and Evening Star by Turns, just as Venus does to us.

Some Astronomers think, they have observed a small Moon belonging to this Planet, about one-fourth as big as Venus: And the Reason we don't frequently see it, is owing, perhaps, to the Unfitness of its Surface to reflect the Light so far. And indeed, if this Planet has no Moon, I can't conceive how the Inhabitants will distinguish their Times, since the Inclination of its Axis is very little +, if any, and the Sun always in their Equator; consequently there must be continually the same Seasons of equal Day and equal Night all round her Orbit.

+ According to Bianchini's Observation, her Axis inclines 75 Degrees, and her Day is equal-

24 of ours,

^{*} Venus and Mercury appearing through the Telescope, sometimes horn'd, and sometimes gibbous, like the Moon, is another Proof of their going round the Sun in Orbits within the Earth's Orbit; and for this Reason they are called inferior Planets. Venus is six Times nearer us at her inferior Conjunction, when on this Side the Sun next us, than at her Superior Conjunction, beyond the Sun. She consequently appears much bigger in the former Situation than in the latter. For though at her inferior Conjunction she shews but a small Part of her enlightened Disk, and looks through the Telescope like a Moon three Days old; yet on Account of her Nearness to us, that small Part contains a greater Area of Light than the whole Disk does, when at her greatest Distance beyond the Sun. In this Situation she has been often seen in the Daynear the Sun, and been taken, by the common People, for a New Star.

Of the EARTH.

HE Earth, on which we live, is one of the Planets; the next beyond Venus. She moves round her Orbit at the Distance of 81 Millions of Miles from the Sun, and completes her Revolution in 365 Days, 5 Hours, and 49 Minutes*. This produces the several Seasons of the Year, Winter, Summer, Autumn, and Spring.

Besides the Annual or yearly Motion, she has a Diurnal one upon her Axis from West to East in 24 Hours, which makes the Sun and all the heavenly Bodies appear to move round from East to West in the same Time. This is the Cause of Day and Night; and of the rising and setting of all the heavenly Bodies.

If the Earth were view'd from the Sun, as she revolves in her Orbit, she would seem to pass (as a small Star) through the Heavens, describing a Circle among the fixed Stars. Which Circle Astronomers have divided into twelve equal Parts, corresponding to the twelve Months of the Year; and have suppos'd each Part cover'd with an Image, or Picture of some living Creature, and call'd that Space by the Name, or Sign of the Image depicted there.

The Names and Characters of the twelve Signs are,

When the Earth, moving round her Orbit, comes opposite to any of these Pictures, she is said to be in that Sign, which the Picture expresses. And as the Earth, like the rest of the Planets, revolves from West to East, the Sun will appear to have an Annual Motion the same Way, and in the same Tract, but still in the opposite Point. For when the Earth is in that Part marked Aries, or the Ram; the Sun will be in that Part mark'd Libra, or the Ballance; and so of any other; as may be seen in the Earth's Orbit delineated on the other Side.——This Circle is called the Ecliptic Line, because the Planets seldom eclipse each other, unless they appear upon that Line, and in the same Point, and then the nearer may happen to obstruct our View of the other beyond it.

* The Revolution of the Earth in her Orbit, which is the Length of a Solar Year, is exactly 365 Days, 5 Hours, and 49 Minutes; but we (to avoid Fractions) account it 365 Days, 6 Hours, which is 11 Minutes too much. These 11 Minutes, in about 134 Years, amount to one whole Day; which Day being retain'd, must make the Sun appear to recede one Day back in the Kalendar in that Time. Confequently, the Vernal Equinox, which happened about the 21st of March, at the Time of the Nicene Council, (in 325,) must, after 134 Years, happen on the 20th, and in 134 Years more, on the 19th, and so on. In our Time, the Equinox was gone back to the 10th of March; i. e. 11 Days from the Place it was in before; and would in Time, have retreated through the whole Kalendar, and thereby have thrown all the moveable Feasts into the greatest Confusion .-- To remedy this Inconvenience, the Legislative Power, by an Act passed in 1752, threw cut the 11 additional Days, by calling the 3d of Septemper the 14th, in Order to bring the Equinox to the Place it was at when that Council was held. And, to keep it fix'd there, order'd, that three Days every 400 Years should be omitted, in the following Manner; viz. the Years 1800 and 1900, which should have been Leap-years, shall be accounted common-Years, of 365 Days only. But the Year 2000, and every fourth Hundred Year after that, shall be a Leap-year of 366 Days, the intermediate Hundreds only common Years. By this Means, our Reckoning will not vary a Day again, in less than 8 or 10000 Years, which is very inconsiderable.

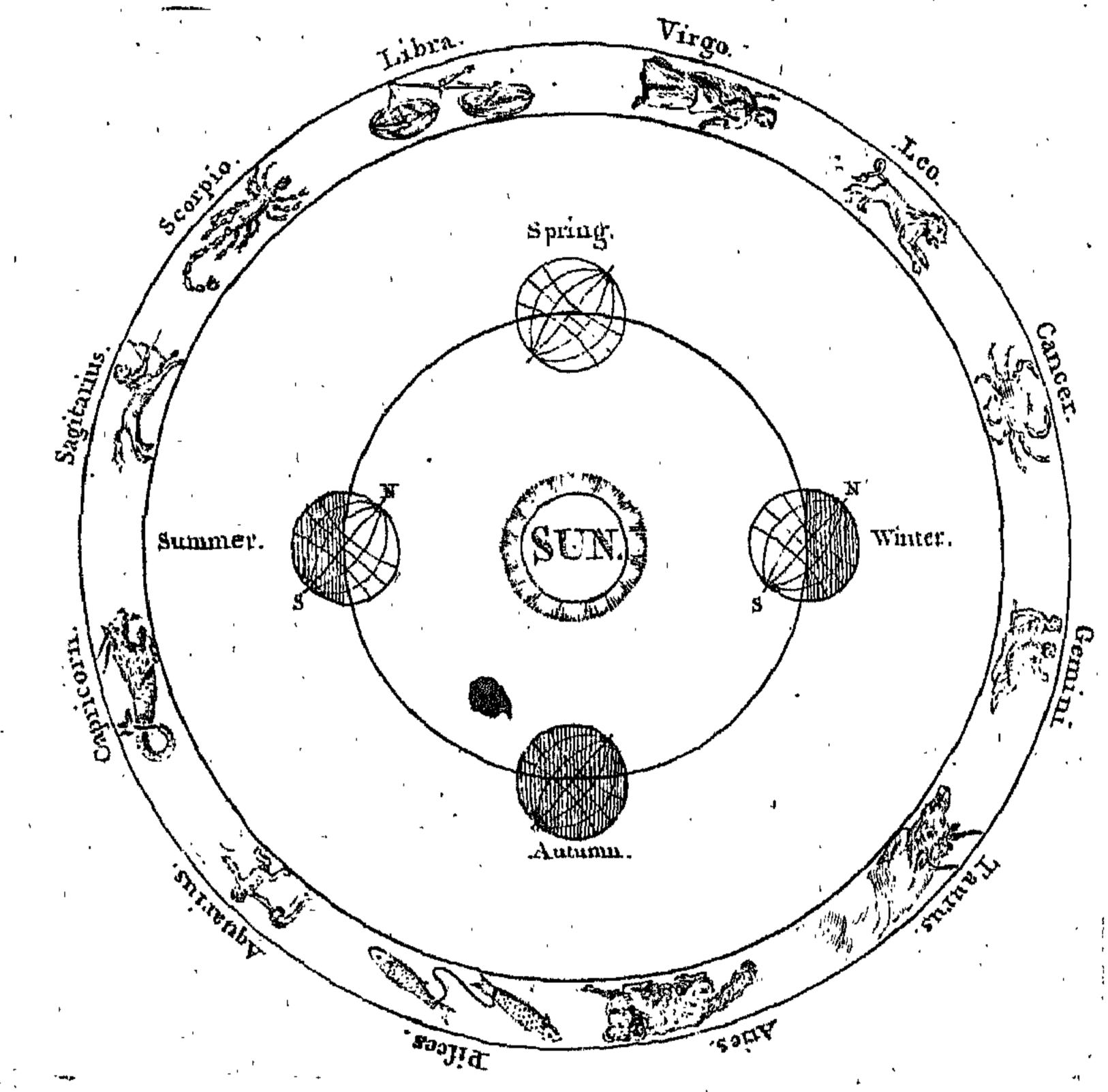
*** ()ur Times and Seasons now correspond with those at the Calling of the first Christian Council, when the Affairs of the Church were settled by Order of the Emperor, Constantine the Great, in the Year of Christ 325.

IF

If the Earth had no Inclination of its Axis to the Plane of its Orbit, the Sun would ever appear in the Equinoctial, making the Days and Nights equal throughout the Year. But by having an Inclination from the Perpendicular of 23½ Degrees, and directing the Poles always to the same Points of the Heavens; hence arises the Division of the Earth into Zones, as exhibited in my Geography, and all the Variety of the Seasons we enjoy; as is clearly expressed in the following Figure.

By this Inclination, the Sun is made sometimes to visit the northern and sometimes the southern Parts, making them more warm and comfortable; and leaving the equinoctial Parts for a small Time, makes them more cool and six

for Habitation.



When the Earth is in Libra, the Sun will appear to be in Aries. Then begins our Spring. The Sun is over the Equinoctial, and his Light is diffused equally from Pole to Pole, making the Days and Nights every where equal. This happens the 21st of March.—When the Earth is arrived at Capricorn, the Sun is then seen at Cancer. Now the North-pole is turned not only towards the Sun, but all the northern Parts are much nearer. His Rays now Atrike us more forcibly; while the southern Parts are turn'd from the Sun; and the Pole itself, and the Parts round it, are involved in Darkness. This is the Beginning of Summer, falls upon June 21st, and the Days are at the longest. From hence, as the Earth advances, the Sun begins to leave the northern Parts, till she arrives at Aries, where the Sun appears in Libra. He is then got over the Equinoctial again, and distributes his Light and Heat equally over the Globe from Pole to Pole, making equal Day and Night, as in the Spring. This is the Beginning of Autumn, and happens September 23d.——As the Earth goes on, the Days still continue to shorten, till she arrives at Cancer, at which Time the Sun will be found in Capricorn. Now the fouthern Parts, you see, are turn'd towards the Sun, and strongly illuminated, making their Summer; whilst the nerthern Parts are turn'd from him, and the Pole is involv'd in Darkness, as the southern was in the opposite Part of the Orbit. Our Days are now at the shortest; the Sun's Rays fall more oblique and feeble; and Winter now begins, December 21st. Here,

Here it may not be improper to observe, that the Orbit of the Earth is not perfectly circular, but a little elliptical: For which Reason, the Sun will be nearer the Earth at one time than another; and the Earth will move sometimes slower and sometimes faster*: Whence it follows, that she will pass over one Half of her Orbit in less time than the other. For, from Autumn September 23d. to Spring March 20th. is eight Days less than from Spring to Autumn again. And so many days is our Summer Half-year longer (when the Sun is farther off) than our Winter Half (when the Sun is nearer) as any one may easily perceive, if he will only count the Days from these Periods in an Almanack.

The Figure of the Earth is that of a round Ball, as is evident from the Observations of those who have sailed round it; as also from its Shadow being always round when it is seen to fall upon the Moon and eclipse her. But the Parts at the Equator have been found, by Sir Isaac Newton, to be higher than those at the Poles by about 17 Miles. And it is necessary it should be so constructed, else the Waters of the Ocean there, would bee whirled over the Land, by the Earth's Diurnal Rotation.

The Circumference of the Earth is supposed to be divided into 360 equal Parts, call'd Degrees, and each of these into 60 Parts, call'd Minutes. Now our Countryman, Mr. Norwood, found by measuring from London to York, in the Year 1655, that one of those Degrees upon the Earth's Surface contain'd 69½ Miles; therefore its Circumference must be 25.020 Miles, and its Diameter 7964.

The Earth is surrounded with a thin vaporous Air, call'd its Atmosphere: This reaches every where to the Height of about 46 or 47 Miles; and serves to suspend the Clouds, surnish us with Winds and Rains, and serves to the common Purposes of Breathing; it is also the Cause of the Morning and Evening Twilight, and all the Brightness and Glory of the Sky.

Every Planet is supposed to be surrounded with such an Atmosphere, which ferves to the same Purposes there, as ours does here f.

* The common People think that the Sun moves; and that the Earth stands fixed in the Centre of the Universe; and when they argue on this Subject, produce those Texts where mention is made of the Foundation and Pillars of the Earth; and that it abideth for ever .- Also, that the Sun rifeth, goeth down, and hasteth to the place where he arose .--- That the Sun cometh forth like a Bridegroom out of his Chamber, and rejoiceth as a Giant to run his Course .-- That the Sun stood still over Gibeon; and went back ten Degrees on Ahaz Dial. But all these Expressions are to be understood according to the Appear ance of Things, not as they dre in Fast: With the Sacred Writers it was usual; when they spake of the Sun or Earth, to express themselves according to the Conceptions and Ideas of the People rather than the real Nature of the System. And even the Astronomers themselves, though they know and teach the Contrary, yet in their common Conversation say—the Sun rises—the Sun sets—the Sun moves; &c. For did they alledge the Motion to be in the Earth, they would be under a necessity of explaining themselves every Time they spake to the Vulgar on that Head.—But there are many Places in Scripture which clearly affert the Situation, Figure, &c. of the Earth, the Plurality of Worlds, Antipodes, &c .- As (1st.) where God is said to frame the Worlds by his Word. Now Worlds must imply more than one, consequently the Planets may be inhabited Worlds .-- He made the round World, i. e. the World, or Earthwe inhabit must be globular .-- He hung the World upon Nothing; i. e. suspended it in the immense Void-He shall subdue People under us; i. e. the Antipodes .-- And Nations under our Feet; i. e. the Kingdoms of the West ... The wind goeth towards the South, and turneth about to the North according to its Circuit; this expresseth the Trade-winds and Monsoons. And the Rev. Mr Kennedy, in his Astronomical Chronology, has proved, from the Writings of Moses, that Time commenced at the Autumnal Equinox -That the Edith began its annial Motion, October 25th at Noon; the Sun being then 10h 24' to the West of Greenwich, i. e. over a Meridian passing through the Middle of the vast Pacific Ocean.-That this happened, in Coincidence with a full Moon, on the fourth Day of the first Week at Noon, being Thursday .-- And from the Year 1761 was 5768 Solar Years.

† If, from what has been said, the Learner understands the various Motions of the Earth, be will easily conceive the Motions of the rest of the Planets. For, they all move in Orbits, like our Earth, round the Sun, which makes, in the same Manner, their Year. They also revolve upon their

Of the Moon.

HE Moon is a Secondary Planet, differing from the rest, because she respects the Earth for the Center of her Motion; and not only her Globe, but her Orbit is carried (as it were) round the Sun together with our Earth; for which Reason she is called the Earth's Satellite, or Attendant.—She makes one Revolution round the Earth, with respect to the fixed Stars, in 27 Days, 7 Hours, and 43 Minutes: This is called the Periodical Month. But the Earth having gone on, in that Time, one twelfth Part of her Orbit, and carried the Moon with her, she will not arrive to the same Position, with regard to the Sun, as she was when she begun her Course, till about $2\frac{1}{2}$ Days more. Hence, from one new Moon to another is about $29\frac{1}{2}$ Days, and this is called the Synodical Month; because, at every new Moon, the Sun, Moon, and Earth are assembled in Conjunction, and lie in a right Line, the Moon being between them.

In every Revolution of the Moon round the Earth, she turns once upon her own Axis, and that is the Reason the same Face is always presented to our View.

The Moon is confiderably smaller than any of the Planets, but by Reason of her Nearness, she appears the biggest; her Distance being only 240.000 Miles from us. Her Diameter is found to be about 2175 Miles;—her Circumserence 6829; which makes her Globe near 50 Times less than ours.

That the Moon has Day and Night, is evident from the Light and Shadow of the Sun seen upon it, moving round it from West to East in one Month.—That she has Summer and Winter is clear from her presenting sometimes more of her northern Parts, and at other times more of her southern; by which also, we learn that her Axis inclines (like our Earth's) to the Plane of her Orbit, but not so much; hers being only 6½ Degrees from the Perpendicular, but ours 23½ Degrees.—That she has Mountains, Seas, and Valleys, is manifest from the Inequalities and Unevennesses in her Surface, which are plainly to be seen with the Telescope.—She is also supposed to have an Atmosphere.—And our Earth doth the Office of a Moon to the Lunarians (if any there) as their Planet does to us.

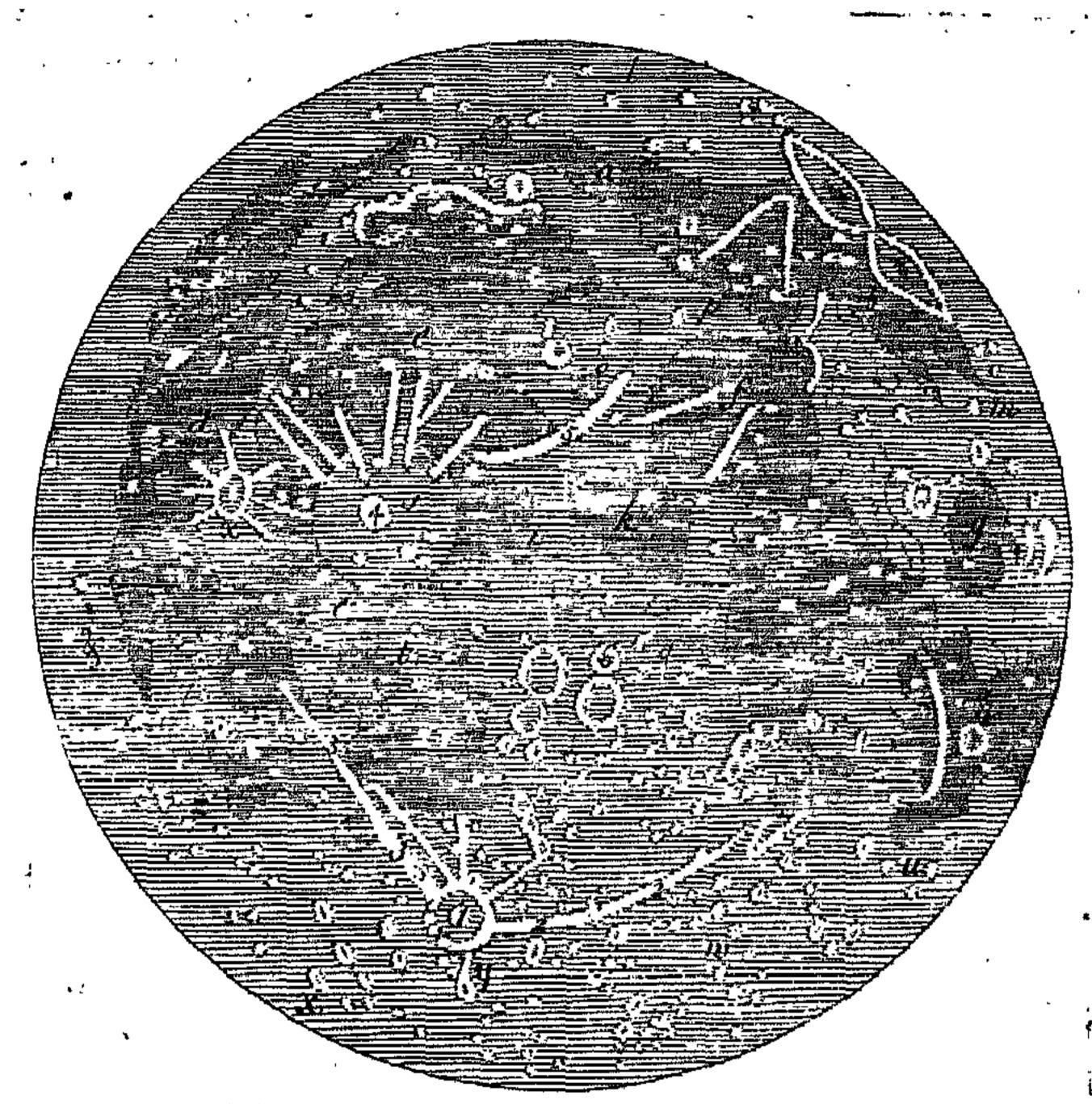
The brighter Parts of the Moon are supposed to be various Regions of Land, shining by the reflected Light of the Sun, the darker Parts are Oceans, Seas, and Lakes, (or deep Cavities and Pits) which, as they absorb the Light, must appear more obscure; the bright Spots and Streaks of Light, are judged to be Islands, Hills and long Ridges of Mountains, strongly illuminated by the Sun.

Axes, which produces the Returns of Day and Night. -- From the Inclination of their Axes to the Plane of their Orbits, arises the Division of their Globes into Torrid, Temperate, and Frigid Zones. -- By keeping their Poles always directed to the same Points of the Heavens, they must necessarily have the varigated Seasons of Summer and Winter. -- And from their Atmospheres arise Twilight, Winds, Miss, &c. as upon our Earth; only, in some Planets, these Phænomena may happen in a greater or less Degree thon upon the Earth itself.

* The Moon appears to move every Day towards the West: This is caused by the Earth's daily Rotation upon her Axis the contrary Way. But her real Motion, like the rest of the Planets, is towards the East. For if you observe her any Night near a fix'd Star, you will find her the Night following considerably remov'd towards the East; and the next Night twice as far; and so on:

Till she (compleating her Round) returns to the Star again.

The Situation of the several Countries, Seas, and Mountains, on that Hemisphere (or Face) which is always presented to our View, is as here delineated



The Occean, Seas, and Lakes are;

- a Mare Hyperboreum d Mare Eoum
- b Paludes Hyperboreæ c Sinus Hyperboreus
- g Palus Meotis
- e Mare Mediterraneum b Mare Caspium f Pontus Euxinus.
- i Mare Adriaticum
- k Propontis.

The Countries and Islands are,

Z	Regio	Hyperborea tia
271	Sarma	tia

n Taurica Chersonsus

o Italia

- p Mæsia Afia
- r Cholchis
- t Peloponnesus u Scythia w Perfia / Sicilia x Arabia
- y Palestina z Ægyptus & Lybia
- à Insula Circunna.

The Mountains and Hills are,

1 Mt. Sinai-2 Mt. Taurus-3 Mt. Sepher-4 Mt. Ætna-5 Mt. Apenninus-6 Mt. Olympus*.

The other Side of the Moon we are never permitted to see, nor the Inhabitants (if any) to see Us, without taking a Journey to this Side next us, and some of them must travel more than 1500 Miles to obtain a Sight of our Earth at all,

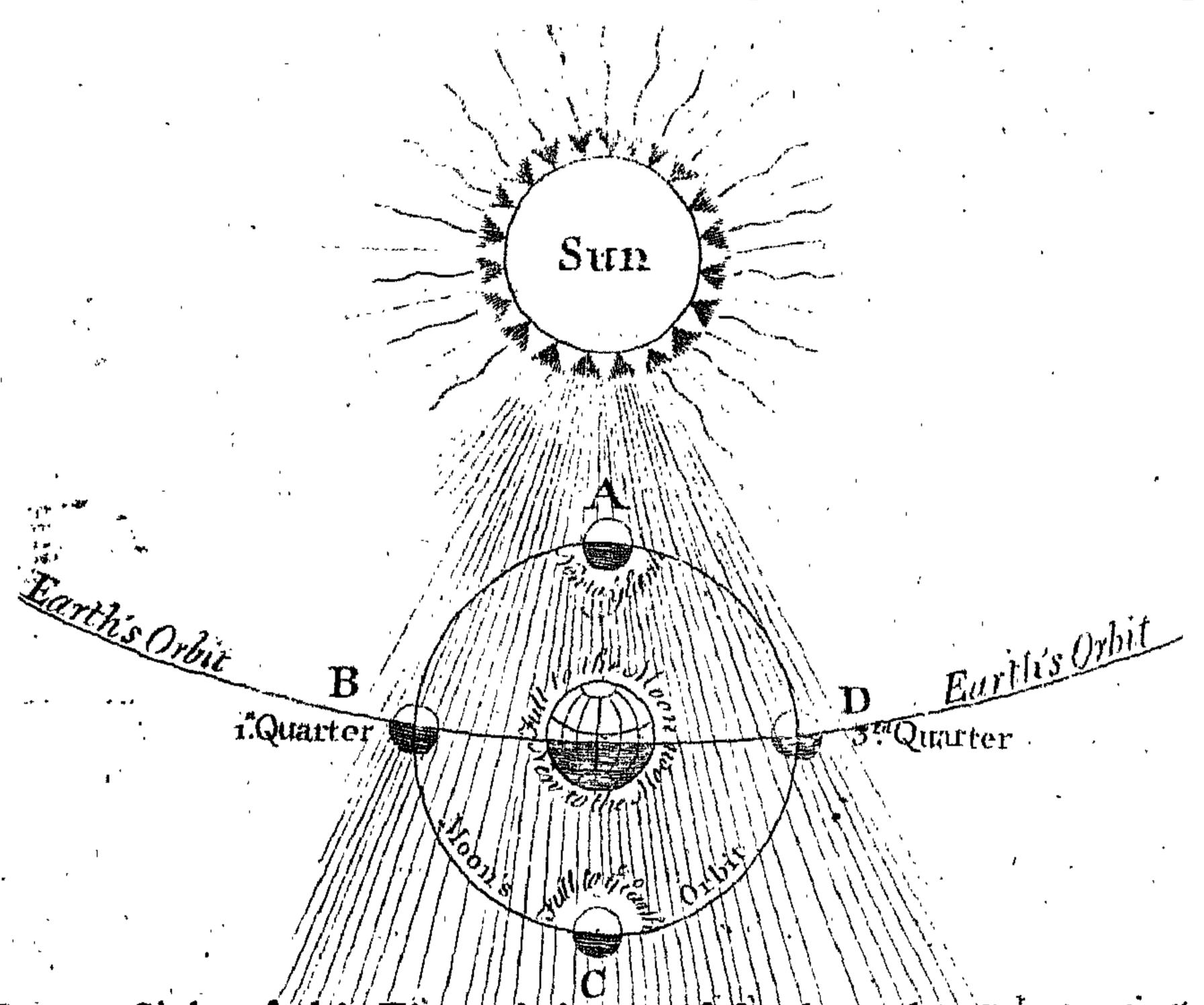
Another Thing, no less remarkable than wonderful and curious, is, that those who inhabit the Middle of the Surface next us (as about Mount Olympus and Mare Adriaticum) will see our Earth constantly over their Heads, and increasing and decreasing in Light like a Moon; whilst those who live near the Borders will see the same Appearances continully in the opposite Parts of their Horizon

To the Inhabitants of the Moon in general, the Magnitude, Light and Heat of the Sun are much the same as with us here, owing to her small Distance from us. But the Length of the Day in that Planet is equal to near 30 of

Some of these Mountains are thought to be higher than any on our Earth. Their Heights are easily obtain'd by Trigonometry. See my Treatile on that Subject. The Shadow of some of the Mountains may be seen with a good Telescope to fall upon the Moon's Surface, and lengthen and shorten as the approaches or recedes from the Sun.

Days and Months are of the same Length; and the Year, to the Inhabitant's there, is exactly equal to ours, because she turns each Pole towards and from the Sun, in the same manner as the Earth does in that Period of Time. But the Variety of Seasons, with regard to Heat in Summer and Cold in Winter is much less than upon this Globe, arising from a much smaller Inclination of her Axis, which is about 6½ Degrees: Consequently the Torrid Zone must be only 13 Degrees—the Frigid Zone 13 Degrees—and the Temperate Zones 77 Degrees broad.

The Lunarians will observe our Earth to shine like a Moon to them, as theirs does to us. For, the Earth and Moon are mutually Moons to one another. Only, when they are a new Moon to us, we are a full Moon to them. And when they appear to increase in Light, we shall appear to decrease in the same Proposition, &c. All this is evident from a bare Inspection of the following Scheme.



From a Sight of this Figure it is manifest, that, that Half of the Earth and Moon next the Sun is always illuminated, and the other Half always dark; Consequently, when the Moon is at A, she is a new Moon, and invisible to us; but we are a full Moon, shining bright to them.—At B, she turns Half her illuminated Side towards us, where she is in the first Quarter; and we appear to the Inhabitants there in the last Quarter.—At C, she becomes a full Moon, shining bright upon us; whilst we are a new Moon, and invisible to them.—At D, she presents Half the illuminated Face again; Now she is in the third Quarter, and we in the first. When she returns to A, her Revolution is complete, and the Appearances begin as before.—As the Moon passes from A round B to C, she gradually increases in Light;—is East of the Sun;—and rises and sets after him.—But in going from C by D to A, she decreases in the same Manner;—is West of the Sun;—and rises and sets before him.

There is one Thing further to be observed, which is, that the Light afforded by the Earth to the Moon is about 15 Times greater (the Earth's Surface being so many Times bigger) than that afforded by the Moon to the Earth: Which strong Resection from the Earth occasions that dusky Light seen on the dark Part of the Moon for some Days near the Change.

Of the Planet Mars.

MARS moves in an Orbit round the Sun, between the Earth and Jupiter, making one Revolution in the Space of 1 Year, 321 Days, 23 Hours, and 27 Minutes; at the Distance of 123 Millions of Miles. This Planet is of a Red Colour, (suppos'd to be owing to the Thickness of its Atmosphere, or Nature of the Soil, to reflect that Redness) and has a Rotation upon his Axis, as appears from the Spots upon his Body, in 24 Hours and 40 Minutes. He is about 4440 Miles in Diameter, which makes him about six Times less in Bulk than our Earth.

Mars, when in Opposition to the Sun, is five Times nearer us than when in Conjunction; for this Reason he appears much larger and brighter at one Time than another. In the Quadrature he appears almost bisseted like the Moon, but never born'd, which is a Proof that his Orbit circumscribes ours, and lies wholly beyond it.

As this Planet is almost Half as far again from the Sun as our Earth is, his Inhabitants will see the Sun's Diameter but little more than Half as big; confequently, his Light and Heat is not Half so great as it is here. Their Year is almost twice as long as ours, and their natural Day is greater by about 40 Minutes.

The Axis of this Planet (as appears from the Revolution of the Spots upon his Surface) is nearly at Right Angles to the Plane of the Orbit; consequently, the Days and Nights are almost equal every where over the whole Globe. But Places in different Latitudes will partake of different Degrees of Heat and Cold, on account of the different Inclination of the Sun's Rays to the Horizon, as in the Case of our Earth, when the Sun is in the Equinoxes, in March and September.

Though Astronomers have not been able to discover any Moon belonging to this Planet, yet it is very probable he has one *, else the Inhabitants will be depriv'd of that most useful Division of Time into Months, which we have measur'd out to us 12 Times in a Year, by that Number of Revolutions of our Moon round the Earth.

The Spectators of the Heavens in Mars will rarely, if ever, fee Mercury, unless it be when that Planet, passing directly between them and the Sun appears as a black Spot upon his Face, as he does sometimes to us. Venus, there, will appear about as far from the Sun as Mercury does with us; and the Earth will appear, as far as Venus appears to us to be from the Sun; and will become their Morning and Evening Star, by Turns, just as Venus does here. And when the Earth, seen from Mars, is in the lowest Part of her Orb near the Sun, she will appear (thro' a Glass) born'd, as Venus does to us in that Position; and her Satellite, the Moon will be of the same Figure; if it can be seen there, and at its greatest Distance not above 4 of a Degree (or a Hand's Breadth) from the Earth. Saturn and Jupiter will appear nearly the same as to us.

^{*} As the Rotation of a Planet upon its Axis makes its Day, and the Revolution round the Orbit its Year; so the Revolution of a Moon about a Planet makes its Month: Consequently, if a Planet has no Moon surrounding it, it can have no Month, but only Days and Years.

Of the Planet Jupit ER.

the Sun in 11 Years, 314 Days, 12 Hours, and 20 Minutes, at the Distance of 424 Millions of Miles from him. By the Telescope, he is found to have dark Shades something like Belts surrounding his Body, near, and parallel to his Equator; in the lowest is a large Spot, by which his Rotation on his Axis was discover'd to be in 9 Hours and 56 Minutes. This Planet, as well as Mars and Saturn is much nearer the Earth when in Opposition to the Sun, than in any other Situation in his Orbit. He is upwards of \$1.000 Miles in Diameter, which makes him about 1000 Times bigger than our Earth; and has 4 Satellites or Moons revolving round him to enlighten him by Night, as our Moon does us, in a constant direct Order, as follows:

Miles

2 revolves
$$\begin{cases} 1 \text{ D. } 18 \text{ H. } 36 \\ 3 - 13 - 15 \end{cases}$$
 at the $\begin{cases} 5^{\frac{2}{3}} \\ 9 \end{cases}$ Semidiameters of $\begin{cases} 229.000 \\ 364.000 \end{cases}$

3 in $\begin{cases} 7 - 3 - 59 \\ 16 - 18 - 30 \end{cases}$ of $\begin{cases} 14^{\frac{1}{2}} \\ 25^{\frac{1}{4}} \end{cases}$ Semidiameters of $\begin{cases} 364.000 \\ 580.000 \\ 1.000.000 \end{cases}$

from his Center.

These Moons always accompany his Body, and never appear to us at the Earth to be further than about $\frac{1}{4}$ of a Degree, or a Hand's Breadth from him: Their Orbits are delineated round Jupiter in the Scheme at the beginning of the Book, but are drawn 100 times too large in Proportion to the Orbits of the Planets. By obferving them with a good Telescope, they may sometimes be seen to cast their Shadows upon his Disk, when they are found between him and the Sun; and at other Times to be eclips'd, by falling into his Shadow as they pass behind him

By observing these *Eclipses* of *Jupiter's Satellites*, it is found that *Light* is not instantly convey'd to us, but takes up 7 or 8 *Minutes* to travel from the *Sun* hither, which is about 8 1 Millions of Miles *.—And this vast Space Light passes through in so short a Time, that the Velocity of its Motion cannot be easily conceived by us.

By the *Eclipses*, or *Immersions* and *Emersions* of these *Satellites*, or *Moons*, into and out of his Shadow, the Longitude of Places on our Globe may be found †.

* When the Earth is between the Sun and Jupiter these Eclipses happen 7 or 8 Minutes too soon; and when the Earth is in the opposite Part of her Orbit, beyond the Sun, they happen as much too late than they should by the best Calculations: The Reason is, because the Light has further to go in the latter Case than the former, by the Diameter of the Earth's Orbit.

[†] Suppose I find, by Mr. White's Ephemeris, that the first Satellite of Jupiter will be immers'd into Jupiter's Shadow at 8 o'Clock at Night, at London, on some certain Day; but, being at Sea, I observe, with a Telescope, the Immersion to begin at 10 o'Clock the same Evening; the Difference, which is 2 Hours turn'd in Degrees, allowing 15° for each Hour, shews that I am 30° to the East, because the Time is more than at London. Had the Time been 2 Hours less, I should have been 30° to the West. By this Method, the Mariners are enabled to correct their Reckoning by these Moons of Jupiter, which were never heard of or seen till discover'd, with the Telescope, by Galileo, on the 7th of January, 1610.

fupiter is little more than 5 Times further from the Sun than the Earth is, consequently his Diameter to that Planet is not a fifth Part of what it appears to us; and, therefore, his Light and Heat must be 25 Times less than ours.

Their Year is almost 12 of ours; but their Days and Nights are but short being about 5 Hours each, the their year is so long and tedious. * As the Axis, of this Planet Inclines but little (if any) to the plane of his Orbit, the Days and Nights must be nearly equal, and the Inhabitants, over the whole Globe, throughout the Year, will enjoy, as they do in Mars, a perpetual Equinox.

The Satellites, or Moons, to the Inhabitants there, will appear as large, perhaps, as our Moon does to us here; and they will have four Kinds of Months, whose Lengths are determined by the Revolutions of their Moons. In one of Jupiter's Years, which is nearly 12 of ours, there will be about 2407 of the least months:—1203 of the fecond Kind:—601 of the third:—and 254 of the fourth or greatest. A Year in Jupiter contains a great Number of their Days, not less than 10464; but of the four Sorts of Months, the least contains only 4 Days;—the fecond about 8;—and the third 17;—and the greatest about 41 Days.

By the frequent *Eclipses* of these four *Moons*; the Navigators in *Jupiter* (if there be any) will be enabled to ascertain their *Longitude* better than we can bere upon our Globe by the Assistance of our (one) *Moon*.

The Astronomers in Jupiter will never see Mercury, Venus, the Earth,—nor perhaps Mars, (unless in their Horizon sometimes, at the Beginning and End of their Twight) since from that Distance, they must appear to accompany the Sun, and rise and set almost at the same Time with him. Nor will they, without better Glasses than ours, be able to know there are such Worlds in Existence. But they will see Saturn and his sive Moons, and perhaps be able to discover the Seasons, &c. in that Planet, something better than we can. Saturn will also appear to them sometimes bigger and sometimes less, as Mars does to us, as he is nearer or further off in his Orbit from them, as is clear from a View of the System itself.

^{*} Jupiter, on Account of his rapid Motion upon his Axis, has his Equatorial Diameter greater than the Polar by about TE Part.

[†] Jupiter will also appear a Moon to them; and to his first Satellite more than so times as large in Diameter as the Earth does to our Moon. He will be seen to increase and decrease in all the Lunar Shapes every 42 ½ Hours.

The Immersions of Jupiter's Satellites into his Shadow can only be seen when he is West of the Sun, and their Emersions when he is East of him. And when he is in exact Opposition to, or Conjunction with the Sun it will be in vain to look for either, because his Shadow, lying perfectly behind his Body, is impossible to be seen at the Earth.—The same Appearances happen with Respect to the Satellites of Saturn.

Of the Planet SATURN.

distant from the Sun 777 Millions of Miles. He makes his Periodical Revolution in 29 Years, 174 Days, and 6 Hours, and 36 Minutes. His Diameter is about 67,860 Miles, therefore he is 600 Times bigger then the Earth, and Half as big as Jupiter. He is observed to have always a round full Face, but is of a dull Lead-colour. It is uncertain whether this Planet revolves upon his Axis or not, because of his vast Distance from us. By viewing him through the Telescope, he appears to have a Ring surrounding his Body, and beyond that 5 Satellites moving round him, as he does round the Sun; and according to the latest Discoveries, their Revolutions and distances from him are as here set down:

From Saturn's Center.

Their Orbits are delineated upon the Orbit of Saturn in the System, but are drawn 100 Times too big in Proportion to the Orbs of the Primary Planets.

Those Moons all move round the Equator of Saturn, and cut his Orbit at an Angle of about 30 or 31 Degrees.

Besides these 5 Moons, this Planet, as I observed before, is sound to have a surprising Ring within them encompassing him on every Side, but doth no where touch his Body. The Breadth of the Ring is about 21.000 Miles, and the Dissance of it from Saturn, on every Part, much the same; so that, in some Situations, the Heavens may be distinctly seen between the Ring and his Body. It appears to be suspended over the Equator of that Planet; is judg'd to be about 7 or 800 Miles thick. And its use is supposed to give Light and Heat to the People there. *

The Ring is nearly parallel to the Earth's Equator, and puts on many different Apearances to Us here: for when Saturn is in 20° of Sagitarius, the nothern Parts of the Planet being turn'd towards the Sun, making it Summer there, the Ring then appears quite open.—When he is at 20° of Pisces, the Ring is quite shut, appearing only as a Line upon the Equator; then their Days and Nights are equal.—When he arrives to 20° of Gemini, the southern Parts are turned towards the Sun, and the Ring appears open again. Now it is Summer to the southern Inhabitants of that Planet.—When he is advanced to 20° of Virgo, the Ring appears shut again, and the Sun being over their Equator, makes their Days and Nights equal, as in the opposite Part of his Orbit.

^{*}This furprising Phenomenon of Saturn's Ring is a modern Discovery; neither were the Satellites of Jupiter or Saturn known to the Antients. The Moons of Jupiter were first discovered, as observed before, by the samous Italian Philosopher Galilæus, with a Telescope, which he first invented in 1610. But Mr. Hugens first discovered the biggest of Saturn's Satellites, which is the fourth, and the Ring, in 1665. The other four Satellites were discovered by Mr. Cassini, the French King's Astronomer, the 3d and 5th in 1671, 1672, and 1673; but the 1st and 2d. were not seen till the Year 1684; with extraordinary Glasses of 100 and 200 Feet in Length.—All Jupiter's Moons—one of Saturn's, the 4th,—and Saturn's Ring, may be seen with a Refracting Telescope of 8 or 10 Feet, or with a Restecting one of 18 Inches or 2 Feet.—With a very good Glass Saturn's Ring appears double, a dark shade going round the Middle of the Broad Part of it.

To the Inhabitants of this Planet, the Sun appears ten Times less than to those on our Earth, as being ten Times further from him than we are; consequently his Light and Heat will be almost 100 Times less *.—Cold and dreary Situation this, when compared to ours! But their Natures no doubt, are suited to it; and our Light and Heat would be as intolerable to Them, as their Cold and Duskiness would be to Us.

Saturn's Year is almost 30 of ours; but the Length of the Days are wholly unknown to us: For we have not been able, on account of his vast Distance from us, to discover his Rotation on his Axis. There is, however, a vast Inequality in the Length of the Days in several Parts of this Planet, and as great Divertity of Summer and Winter and other Seasons, owing to the Inclination of the Equator to the Plane of Saturn's Orbit, which is about 30 or 31 Degrees.—Hence to the Inhabitants in the Latitude of 60 Degrees, their longest Day will have no Night, and, on the contrary, the longest Night will have no Day, just as with us under the Artic Circle.

The Saturnians (without better Optics than ours) see none of the Planets but Jupiter, and he appears to accompany the Sun, being never found, either before or after him, more than 34 or 35 Degrees. He therefore becomes their Morning and Evening Star by Turns as Venus does to us. But they will see their own five Moons, perhaps, larger than ours, which will measure out to them five Kinds of Months of different Lengths. These Moons will increase and decrease, come to the new and full, and frequently eclipse each other and the Sun (as ours does here) as they revolve in their several Orbits round him. He likewise does the Office of a larger Moon to them, waxing and wancing to each as our Earth does to our Moon.

The amazing Ring suspended in their Sky, will cast a Shadow (as the Sun moves Northward or Southward of their Equator) over vast Regions of Saturn's Body, which removing from one Part to another, will cause great Changes in the Light and Darkness on that Planet. And it is very probable the Inhabitants may be ignorant of the Reason of these wonderful Varieties. For though we, here at the Earth, stand convenient to see it, and know it is owing to a mighty Ring surrounding his Body, yet it is not easy for them there to discover it. They must naturally imagine it to be in the Heavens, and have no Way of determining the Distance but from Observations made in different Latitudes. For tho' at their Equator it appears over their Heads, yet as they recede either Northwards or Southwards, the Ring will appear lower till at the Latitude of so, perhaps, it will not be visible. So that from thence to their Poles, the Inhabitants (if any) will be totally ignorant of that wonderful Phænomenon. Nor can they see it without taking a Journey on purpose.

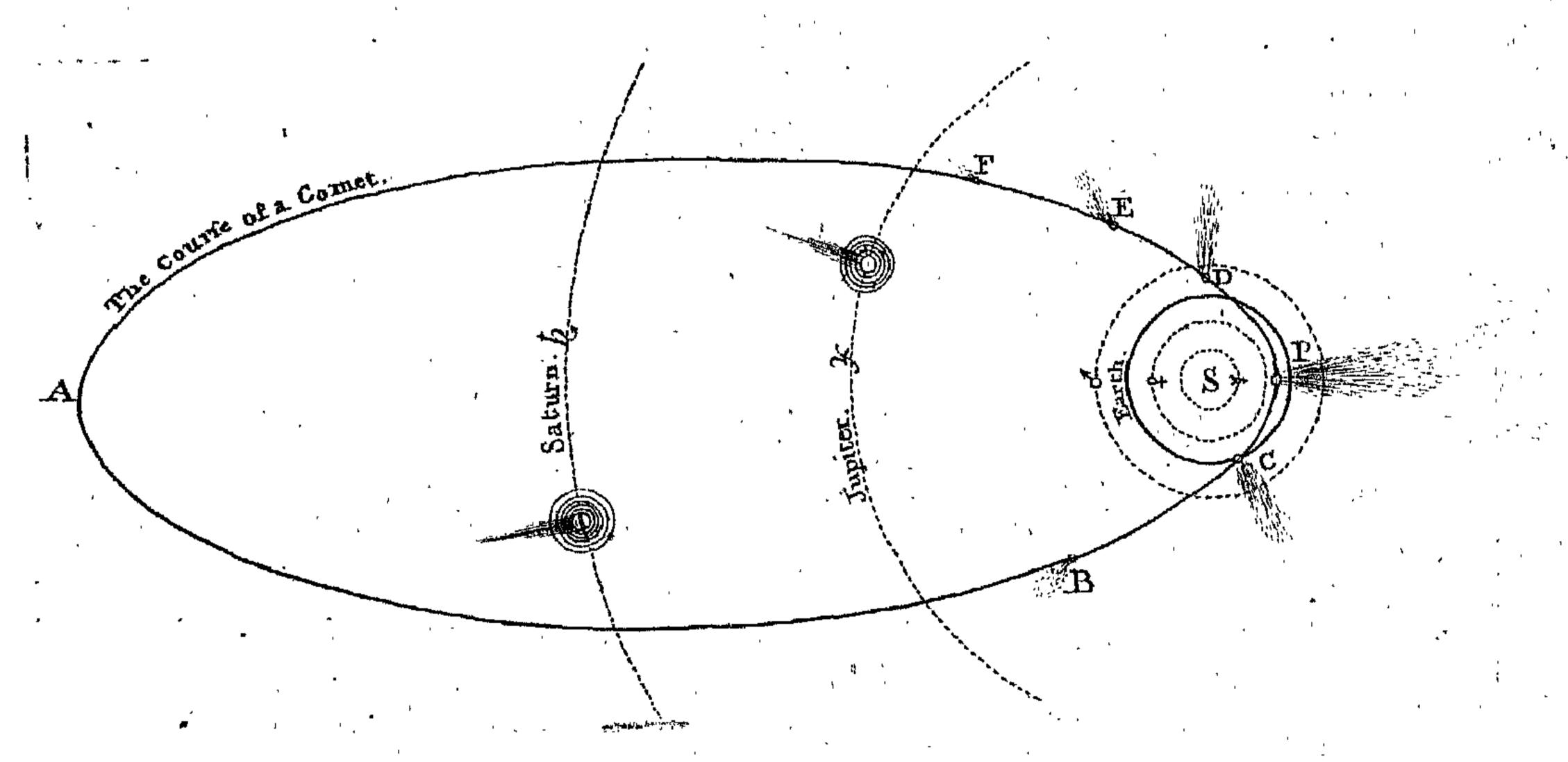
The Diameters of the heavenly Bodies appear to increase and decrease, as we approach or recede from them, directly as the Distance themselves; i.e. at twice the Distance they appear Half as large; at thrice the distance three Times less-But the Proportions of Light and Heat increase and decrease inversely, as the Squares of those distances; i.e. at twice the Distance it is four Times less light and hot; at three Times the Distance, nine Times less-And the Bulk, or solid Contents of those Bodies are to each other directly as the Cubes of their Diameters; i.e. if the Diameter of one Planet be don-ble the Diameter of another, then the Bulk of the larger is eight Times greater than that of the smaller; if the Diameter be three Times greater, then it is twenty-seven Times bigger; and so of any other.

Of the Comers.

ESIDES the Planets already treated of, there are other large Bodies belonging to our System, which are (as it were) a Kind of temporary Planets; for they sometimes make their Appearance in our Regions for a while, and then disappear for a certain Space of Time, after which they return again.—These wandering Bodies are call'd Comets or Blazing Stars.

The Antients thought them to be Meteors or Exhalations, set on fire in the middle Regions of the Air: But the modern Astronomers have found that they are large globular Bodies, moving in various Directions across the System; and that their Orbits are not circular, like those of the Planets, but very eliptical, or oval; and therefore, they are sometimes at a moderate Distance from us; at other Times, they ascend to vast Heights above Saturn, and so become invisible, till they return into our Part of the Heavens again.

The Manner in which a Comet revolves in its Orbit round the Sun, and through the System is here delineated.



In this Scheme, S represents the Sun; A, B, C, P, D, E, F, the ecliptic Orbit of a Comet; the black Circle C, P, D, the Orbit of the Earth; and the dotted Circles the Orbits of the rest of the Planets. Now the Comets are sound to revolve round the Sun in the same Manner as the Planets, differing from them only in the Form of their Orbits; for in the Planets their Orbs are nearly circular, but those of the Comets are long Elipses. And the greater the Elipses are, the surther will they stray from the Sun at one Time, and the nearer approach him at another: consequently, they must suffer great and different Degrees of Heat and Cold. When a Comet is at A, it is at its greatest Distance from the Sun, sar beyond the Orbits of the most distant Planets; and, as they are not much larger than the Moon, (some sew, perhaps, as large as the Earth) they remain hid from our Sight. But as they descend within the System, they become visible, appearing, at first, like Stars of the smallest Magnitude; and as they approach the Sun they continually grow larger, till, at last, they exceed Stars of the largest Size.

When the Comet arrives at or about B, near the Orbit of Mars, it will be seen, perhaps, beginning to emit a small lucid fiery Tail, from the Heat it then receives from the Sun. As it approaches nearer, the Heat increasing, the Tail grows longer. At the Peribelion, at P, and a little beyond it, the Tail is longest of all. From thence, as the Comet continually ascends, and the Heat diminishes, the fiery Tail begins to contract its Dimensions, and grow less and less, till it becomes invisible about F; after which it is seen no more, without a Telescope, till it returns again.

The Comets themselves are suppos'd to be hard, dense, durable Substances, like the rest of the Planets; but their Tails seem to be owing to some peculiar, rare and luminous Matter, in their Atmosphere, which is easily excited by the Sun's Heat, and thrown off in siery Vapours, always into the Regions opposite the Sun *. The Magnitude and Length of the Tail, depend chiefly on the Heat the Comet receives from the Sun; whence, the nearer it approaches the Sun, the longer is its Tail.

The Tails of the Comets are observed to put on various Appearances. For sometimes they appear of a vast Length, growing wider and thinner, the farther they proceed from the Body of the Comet. At other times they appear shorter, and with a narrower Termination, like a Sword. And sometimes, the Tail seems to surround the Comet, with fine lucid Beams, like Hair. All which Variety of Aspect arises from the Position of the Tail with respect to us at the Earth.

Some of the Tails, near their Extremities, are so very fine and transparent that the fix'd Stars may be distinctly seen through them: And their Lengths become sometimes so amazing, as to take up more than 40 Degrees in the Heavens; which, considering their Distances at that Time, cannot measure less that 70 or 80 Millions of Miles.

The Comets moving in very long eliptical Orbits, and becoming visible to us only through the small Part of their Course near the Sun, and remaining lost, as it were, and unknown to us through their immense Journies beyond Saturn, their Periods cannot be so easily determined as those of the Planets. And when a Comet does return, we are not certain it is the same, but by comparing its Direction through the Heavens with some one that has appeared there before.

By this Method, we know that the Period of the Comet which appear'd in 1680 is 575 Years, because one of the same Kind was seen in the Year 1106; and once before about the Year 532; and also 44 Years before Christ. This Comet, at the time it appear'd last, came so near the Sun, as to be within one Sixth Part of the Sun's Diameter from the Sun's Body; consequently, it must receive a Degree of Heat 28.000 Times hotter than our Earth in Summer; which is about 2000 Times hotter than red-bot Iron.—This is a Proof that the Bodies of the Comets are extremely fix'd and durable, else they must be totally dissipated and consum'd by such an intense Heat.

^{*} As to the Cause of the Ascent of the Tail, Kepler ascribes it to the Rarefaction of the Comet's Atmosphere by the Heat of the Sun, and the impulsive Force of the Sun's Beams, carrying along with it the Matter of the Comet's Tail; which also accounts, at the same Time, for the Direction or Position of the Tail, which is always toward the Parts opposite to the Sun.

The Revolution of another Comet, which appear'd in 1682, is suppos'd to be 75½ Years, because such an one was seen in 1607, 1531, and 1456. Its Return was expected in 1758, but was not seen till the Beginning of 1759, at which Time it was visible in many Parts of America; and observed to be ascending back from the Center of the System. This Comet, at its nearest Approach to the Sun came within the Orbit of Venus, and receiv'd a Degree of Heat almost equal to that of boiling Water.

Another Comet, which was seen in 1661, and before in 1532, will probably return again in the Year 1789, making its Tour round the Sun in 129 Years.

The Number of the Comets is imagin'd to be about 24 or 25; but these Periods are all that are certainly known to our Astronomers at present; and are what we must content ourselves with, till surther Observations and Discoveries are made.

These amazing Bodies, in their Journies to and from the Sun, move in all Manner of Directions across the Orbits of the Planets*. For some revolve from West to East; others, on the contrary, from East to West: Some, again, move from North to South; and others nearly perpendicular to the Plane of the Planets Orbits. And as they are sometimes burning with intense Heat in the Neighbourhood of the Sun, and at other Times scarcely receiving any friendly Influences from him; it cannot be supposed that they are Places of Habitation; and design d for Animals and Vegetables, as the Planets are.

What then may be the Uses of these wonderful Bodies and their fiery Tails to the System, is, at present, only conjectural.——Some have supposed, that in each Revolution they accede nearer the Sun, till at last, falling upon his Surface, they supply him with fresh Fuel.—Others have thought, that their Tails are fine volatile Spirits excited by the Heat of the Sun and dissipated through the System, to refresh the Atmospheres of the Planets, and supply them with those vivisying Spirits so necessary to sustain the Life of Animals and Vegetables there.—Some, again, have look'd on them, on Account of the excessive Heat they receive near the Sun, and the intense Cold they must suffer at their greatest Distance from him, to be proper Receptions for the Habitation of the Damn'd.—Whilst others have concluded them, with greater Probability, to be the Executioners of God's Vengeance on finful Worlds; by scattering their baneful Influences on the Inhabitants, or dashing the Planet to Pieces, and reducing it to its chaotic State again.

The learned Mr. Whiston has made it very probable, that a Comet passing by our Earth, in its Descent towards the Sun, and involving us in its Tail, in the Days of Noah, was the Cause of the Deluge; and that another, sometime hence, in its Ascent from the Sun, after having been thoroughly ignited there, may bring about the General Constagration.

* As many of the Comets are observed to revolve in a Direction contrary to the Planets, and with a great and regular Velocity; we are assured there can be no folid Orbs, as the Antients supposed; nor Vortices, or Whirlpools, of subtle Matter in the Heavens to carry the Planets round the Sun, as the Cartesians imagined. For if there were such Circulations of etherial or subtle Matter, the Comets would, when they entered into the Regions of the Planets, be necessarily driven from their Course by the rapid Motion of such a mighty Torrent, and be carried the same Way with the neighbouring Bodies. But, as ocular Demonstration assures us that no such Thing happens to the Comets, but that they preserve their Motions with the great est Freedom, as in a perfect Void or Vacuum; we must acknowledge, that in the Heavens there is no Resistance, and consequently, no Medium, or Fluid, which can have any sensible Density; but that all the Spaces between and beyond the Planetary System are an immense Vacuity, admitting nothing but the Rays of Light, unless near the Atmospheres of the Comets and Planets; and, consequently, that all Solutions of the Phænomena of the Heavens, depending on the Supposition of such a subtle Matter, are sale, and contrary to the evident State of our System.

Of the Fix'd Stars.

Ta vast Distance beyond the Orbits of all the Planets and Comets, we behold, in a clear Night, an azure Concave studded over with Stars of different Lustres and Magnitudes. And because they retain the same Distances and Situation with respect to each other at all Times, they are justly call'd fix'd Stars, to distinguish them from the Planets, which are continually wandering from Place to Place. Their Distance from us must be immensely great, since there is no visible Alteration in their Positions, with respect to each other, when viewed from different Parts of the Earth's Orbit; consequently, the whole Orbit of the Earth, (which is 162 Millions of Miles in Diameter,) is but a Point in Comparison of it.

Astronomers have computed the Distance of the greatest, and consequently the nearest fix'd Star, to be about Two Million of Millions of Miles *; which is so very great, that a Cannon-ball would scarcely arrive thither in 700,000 Years, though it should roll on with the same Velocity it receiv'd at the Mouth of the Cannon. And it is very probable that they are situate as far from one another.

From the fix'd Stars being visible at such an immense Distance, we conclude that they do not shine with a reflected or borrow'd Light, like the Planets which are near us, but with their own native Glory, which they must have in themselves, like our Sun; and, therefore, are bodies similiar to him. And were we to be remov'd as near the fix'd Stars as we are to our Sun, they would appear as large, perhaps, as he does; and our Sun, beheld from thence, would appear a twinkling Star among the Rest.

The Stars, as they appear of different Sizes, are divided, for Distinction's sake, into Six different Magnitudes or Classes; of which, the largest and brightest are call'd of the first Magnitude. Next these, are the second Magnitude; and the next Size, are the Stars of the third Magnitude; and so on to the sixth, which are the smallest that can be seen by the naked Eye in the clearest Night; but it is seldom that we can see those that are smaller than the fourth Magnitude.

Several Catalogues have been made of the fix'd Stars: The first was by Hipparchus, the Rhodian, on seeing a New Star ‡ in the Heavens, about 120 Years before Christ; his Catalogue contain'd 1022 Stars. After him, Ptolomy enlarg'd that Catalogue to 1026. Ulug Beigh, the Grandson of Tamerlane the Great, made a Catalogue of 1017 Stars. Tycho Brahe determin'd the Places

* Mr. Hugens supposing the fix'd Stars of the same Magnitude with our Sun, found that Syrius, or the Great Dog, which is the largest, and therefore, very propably, the nearest, appear'd about 27.000 Times less than the Sun: His Distance must therefore be 27.000 Times as far; that is, upwards of 2 Millions of Millions of Miles.—Dr. Bradley, from a very subtle Calculation makes the Distance 10 times greater. According to him, the bright Star in the Head of Drace is 400.000 times further off than the Sun.

+ The Twinkling of the fix'd Stars is owing to the exceeding Smallness of their apparent Diameters occasion'd by their immense Distance; so that every little Particle of Dust that floats in the Air, when it comes between the Star and the Eye will eclipse it; and as the Air is full of Various Kinds of Particles, some of them are constantly passing between the Eye and the Star, and consequently cause the Star to twinkle.

‡ It is seldom we are favour'd with the Appearance of a New Star; but at present there is a wonderful one in the Neck of the Whale, which is observed to appear and disappear periodically, its Period being Seven Revolutions in Six Years.

of 777 fix'd Stars, and reduc'd them to a Catalogue. Kepler's Catalogue contain'd 1163. The Prince of Hesse's Catalogue was of 400 Stars. Ricciolus enlarg'd Kepler's Catalogue to 1468. Bayerus is said to have describ'd the Places of 1725, and mark'd each Star with a Letter of the Greek Alphabet: The biggest Star in each Constellation being denoted by the first Letter; the next Size by the second; and so on to the smallest. After him, Hevelius*, of Dantzick, compos'd a new Catalogue of 1888. But the largest and best is that of Mr. Flamstead's, which contains about 3000; of which Number it is seldom that a good Eye can see more than 100 together: And Mr. Flamstead himself afferts, that the naked Eye can discover no more than 384 Stars, in the clearest Night, in both Hemispheres. But, with an ordinary Telescope, we can discover in some Parts of the Heavens, 10 Times as many as were visible to the Eye before; so that we have Reason to believe, that only the all-wise and powerful Being, who fram'd them, is able to tell the real Number of the Stars, and call them by their Names.

The ancient Astronomers, the better to describe and distinguish the fix'd Stars from one another, have drawn them upon Globes and Maps, in 48 Images or Parcels as of Men, Lions, Bears, Triangles, Crowns, &c. whence it comes to pass, that every Star has a Name from that Part of the Image it is situated in; as, the Bull's Eye, Lion's Heart, Orion's Belt, &c.—Some modern Astronomers have added 30 Constellations more, which were unknown to the Antients. Some of which are form'd from the Stars about the South Pole, and others from the Stars lying between the old Constellations: these Latter are depicted on the Globe in a fainter Character.

The Names of the several Constellations, and the Number of Stars visible in each, you will find upon the Cælestial Globe, laid down exactly corresponding with those in the Heavens.

The Constellations on the North Side of the Ecliptic or Zodiac, are 34.

The Little Bear, Great Bear, Boötes, Crown, Dragon, Cæphus, Hercules, Harp, Swan, Perseus, Andromeda, Cassiopeia, Auriga, Lynz, Little Lion, Greyhounds, Charles's Heart, Lizard, Serpent-holder, Serpent, Dart, Camelopardus, Berenice's Hair, Mount Menelaus, Eagle, Goose, Fox, Antinoüs, Dolphin, Little Horse, Pegasus, Great Triangle, Little Triangle, and Fly.

The 12 Signs, or Constellations, of the Ecliptic, or Zodiac, through which the Planets all move in their yearly Revolutions, are,

** Ram, & Bull, II Twins, ECrab, & Lion, IN Virgin, Scales, In Scorpion, Archer, & Goat, Water-bearer, & Fishes.

The Constellations on the South Side of the Ecliptic or Zodiac, are 32.

The Whale, Orion, River Po, Hare, Great Dog, Little Dog, Ship, Crow, Hydra, Sobiefki'-Shield, Cup, Centaur, Wolf, Altar, South Crown, South Fish, Phanix, Crane, Unicorn, Indian, Peacock, Dove, Chamaelion, Bird of Paradise, Indian Fly, Flying Fish, Sword Fish, South Triangle, Oak, Toucan, Hydrus, and the Cross.

Hevelius made to new Constellations out of the unform'd Stars lying between the other Con-stellations; they are express'd on the Globe in a sainter Drawing.

Befides these Constellations, there is a whitish Trast in the Heavens, call'd the Milky Way, passing round thro' Auriga, southward, by the Twins, the Great Dog, the Ship, and returns by the Scorpion, Archer, Serpent-bearer, Eagle, Swan, Cassiopeia, and Perseus, back to Auriga again. This Path is of unequal Breadth, and in some Parts, dividing itself, becomes double. It is suppos'd to be form'd by an infinite Number of small Stars, visible only through the Telescope, which combining their Light in those Parts, cause that shining Whiteness. We have a fine View of it, in the Evenings, in the Months of February and August. There are several dusky or cloudy Spots in the Heavens, (particularly one in the Crab, another in the Middle of Orion's Sword, &c.) which are suppos'd to be owing to the same Cause as the Milky Way*.

Some have wonder'd at the great Irregularity which appears in the Disposition of the fix'd Stars; but that Irregularity is of the greatest Use: For, from thence it is the Heavens become divisible into Signs and Constellations, which a regular Situation would have render'd impossible. By this means we are enabled to ascertain, at all Times, the Revolutions and Places of the Planets, and to render their Motions serviceable to us here.

The rising and setting of the fix'd Stars, their apparent Motions every Night to the West, is owing to the Earth's Rotation upon its Axis, every Day, toward the East. And our seeing different Stars, at different Seasons of the Year, is owing to the Earth's Revolution in its Orbit round the Sun, as delineated at Page 8.

The whole Body of the fix'd Stars appears to have a flow progressive Motion in (parallel to the Ecliptic) towards the East, of about 50 Seconds yearly, i. e. 1 Degree in 72 Years; consequently, they do not complete one Revolution in less than 25.920 Years; after which Time they all return to the same Places again. This long Period was call'd the Great Year, and the Antients imagin'd that when it was finish'd all Things would begin anew, and return in the same Order and Manner as before.

† This Motion of the fixed Stars is not real (like that in the Planets) but apparent only; being caus'd by the spheroidical Figure of the Earth, turning its Axis continually, by little and little, (from its Parallelism) towards the West; which, consequently, must make all the Stars appear to advance forward just as much towards the East.

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^{*} Some of the largest Stars have particular Names given them; as Aldebaran, a large Star in the Bull's Eye; --- Castor and Pollux, two remarkable Stars in the Twins; --- Regulus and Deneb, both of the first Magnitude, the former in the Lion's Heart, the other in his Tail--- Syrius, a glorious Star in the Jaws of the Great Dog; --- with many others, which you will find inserted in their proper Place on the celestial Globe. The Ancients placed these Figures in the Heavens, either to commemorate the Deeds of some great Men, or some remarkable Exploit or Action; or else took them from the Fables of their Religion, &c. And the Moderns retain them, to avoid the Consustion which would arise by making new ones, when they compare the modern Observations with the old ones. -- Those Stars which lie between the several Constellations, not cover'd by any Figure, are call'd unform'd Stars. The Division of the Stars into Images and Figures is of great Antiquity, probably as old as Astronomy itself. For in the most antient Book of Job we read of the Constellations Orion, Arcturus, the Pleiades, and Mazzaroth; which last is supposed to express the 12 Signs of the Zodiac.

Of the Immensity of the Universe.

HE Ancients look'd upon the azure Firmament of the fix'd Stars to be the Boundary of the Universe, and that they were plac'd (as it were) in a solid Arch, at an equal Distance from us. But the Moderns, since the Invention of Telescopes, have suppos'd, and with greater Probability, that every fix'd Star is a Sun surrounded with Planets and Moons like our own. Which several Systems of the fix'd Stars, as they are at great and sufficient Distances from the Sun and Us, so are they conceiv'd to be at proper Distances from one another. Hence it is that the fix'd Stars, though perhaps nearly of the same Size, appear to us of different Magnitudes;—the nearest, largest, those further off, less and less.

To suppose these glorious Bodies were all created only for our Use;—purely to scatter over our Globe a dim Light in a Winter's Night, is forming very weak I-deas of the Divine Wisdom; since one Luminary deposited much nearer, would have been more beneficial to us, than all their twinkling Rays now united together.—Far more rational, as well as more to the Glory of God, is the Supposition that every fix'd Star is a Sun, surrounded with a System of Planets, which at different Distances and in different Periods of Time perform their diurnal and annual Revolutions round him.

In what Part of the Universe our System is plac'd;—whether in the Center; or nearer the Extremities;—or whether there are not nobler Systems somewhere through the amazing Workmanship of God, are Problems beyond our Abilities at present to determine *. For by Reason of their vast Distances, our best Telescopes are found too weak to discover actual Globes revolving round their respective Suns, yet we can clearly enough discover that the fix'd Stars have in themselves the Nature of Suns; and that some Things, similar to Planets, sometimes appear and disappear in the Regions or Neighbourhood of those Bodies.

What an amazing Scene does this display to us!—What an inconceivable Vastness and Magnificence of Power and Wisdom does such a Frame unfold!—Suns beyond Suns, to our weak Sense indefinitely distant from each other, and Myriads of Beings dissus'd thro' and peopling Infinity, all subject to the Creators Will!—An Universe of Worlds, deck'd with Mountains, Lakes, and Seas; with Rivers, Trees, and Animals, various as the Globes they inhabit!—All the Produce of indulgent Wisdom to cheer Infinity with endless Beings, to whom his Omnipotence may give a variegated eternal Life.

The Rev. Mr. Hervey, in his Meditations, has given a most magnificent Description of the Universe. Could we (says that Author) wing our Way to the highest apparent Star, we should there is see other Skies expanded, other Suns that distribute their inexhaustable Beams of Day, other stars that gild the alternate Night, and other, perhaps nobler Systems established; established in unknown Prosusion through the boundless Dimensions of Space. Nor do the Dominions of the great Sovereign end there: Even at the End of this vast Tour, we should find ourselves advanced no further than the Frontiers of Creation, arrived only at the Suburbs of the great Jehovah's Kingdom."

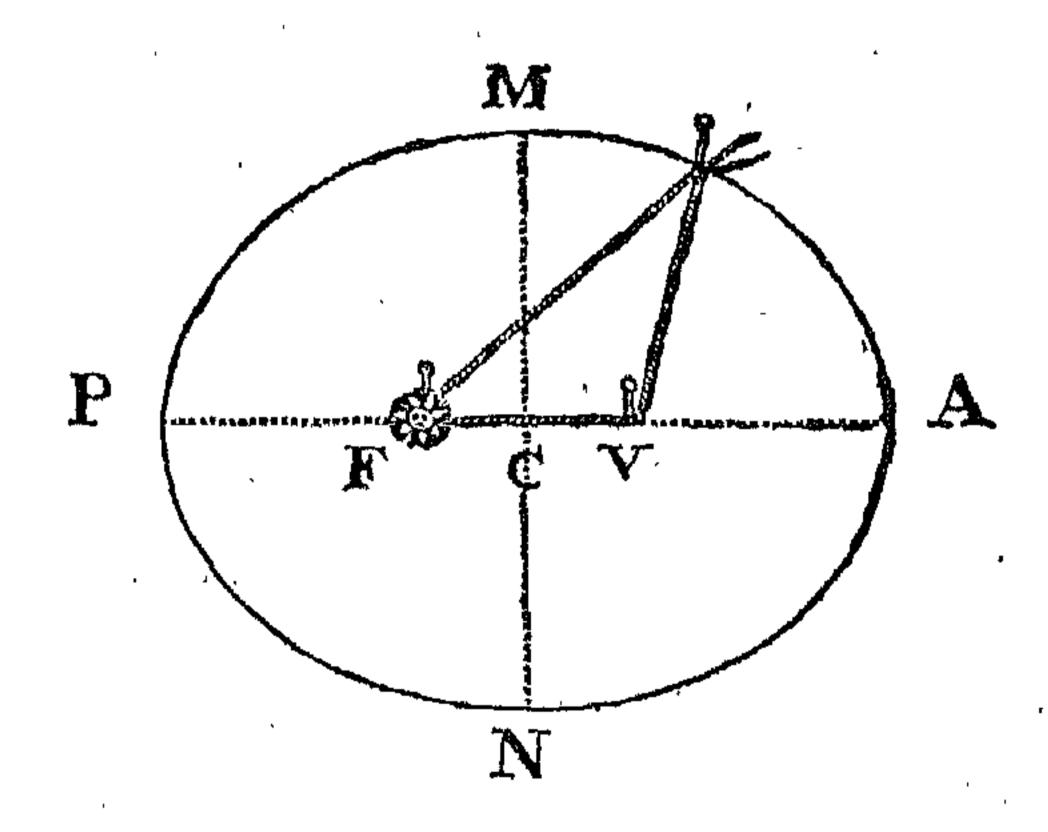
Of the true Figure of the Planets Orbits:

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Nature of their Motions in them; and their Situations with refpect to the Ecliptic; also their Aphelions, Perihelions, Eccentricities, Inclinations, and Nodes.

HE Planets do not revolve round the Sun in Orbits exactly circular, but a little elliptical; consequently, they are sometimes a little nearer, and sometimes a little further from him. For the better understanding their Motion, it will be necessary to shew how to delineate an Ellipsis; which the Learner may take as follows.

Fix upright (upon any Plane) two Pins, as at F and V; round these tie a Thread, something longer than their Distance from each other; then apply, in the Doubling of the Thread, another Pin, or Pencil, so as to keep the Thread properly strain'd; and in that Manner carrying the Pin round, with a steady Hand, the Point of it will describe a Curve we call an Ellipsis.—The nearer the Pins are together the more circular will the Figure be; so that when they unite, or stand together, the Thread will describe a perfect



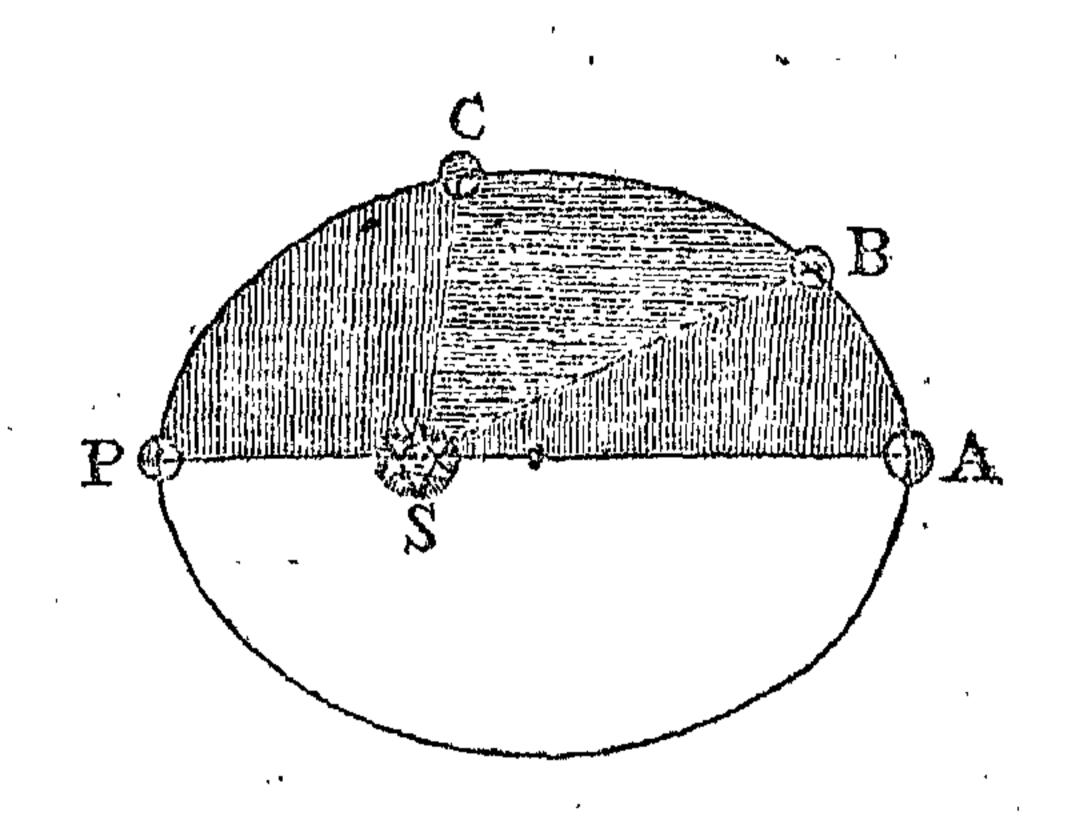
Circle.—The two Points, F and V, where the Pins stand, are call'd the two Focusses of the Ellipsis. The Line PA, going through the Focusses, from one End to the other, is called the longer Diameter; and the Line MN, cutting the former in the Middle between the Focusses, is the shorter Diameter. The Point where these two Diameters intersect each other at C, is the Center of the Ellipsis.

Now each of the primary Planets in their yearly Revolutions round the Sun, deferibes a Line of this kind : For the Sun is not plac'd in the Middle of their Orbs, at C, but in one of the Focusses.——Suppose the Sun to be plac'd at F; then will the Planet, when at P, be nearest the Sun, and at A furthest from him. The Point P is call'd the Peribelion of a Planet; and A the Aphelion: And the Points M and N are the middle or mean Distances, because the Distances FM and FN are a mean between FP the least, and FA the greatest Distance. The Distance FC, or CV, is call'd the Eccentricity, which is different in the different Orbits of the Planets, but in all of them it is so little, that in small Schemes made to represent their Orbits, it is needless (and almost impossible) to express it. But, in the Orbits of the Comets, the Eccentricity is very considerable; for their Orbs are very long Ellipses, and the Focusses at vast Distances from one another.

^{*} The Way we came to know that the Orbit of the Earth is elliptical, is, because the Sun appears of different Sizes at different Times of the Year. In June his Diameter is about 313 Minutes; but in December, 323 Minutes. The Earth also is found to move flower in the former Case than in the latter. It must therefore, necessarily follow, that we are further from the Sun in Summer than in Winter, and that our Orbit must be elliptical.

As the Planets are sometimes nearer, and sometimes further off the Sun, the Velocity of their Motions must be different in different Parts of their Orbits.—For, when the Planet is at A, in its Aphelion, (or most distant Part of its Orbit) the Sun's Action upon it will be less than in any other Part; consequently, its Motion there will be the flowest. But, as the Planet proceeds from A towards P, in the Figure below, its Motion (as it is continually coming nearer the Sun) will be continually accelerating, or increasing, till it arrives at P; where the Sun acting upon it with the greatest Force, the Planet will move with the greatest Swiftness. And as it revolves from P, its Peribelion, it will slacken, by little and little, its Pace, as it recedes surther from the Sun, till it comes to A, where the Motion is flowest of all.—Now, since the Motion of the Planets is so unequal, the Arches thro' which they pass in certain given Times, must be so too.

Though the Planets move through unequal Spaces in equal Times, yet it has been discover'd, that the triangular Spaces within the Orbit, made by Rays drawn from the Sun to the Planet will be always equal in equal Times.—For, if the Times in which the Planet moves from A to B, and from B to C, and from C to P, be equal to each other, as, suppose, a Day, a Week, a Month, &c. the Areas, or triangular Spaces, ASB, BSC, and CSP, describ'd in those equal Times, (by



Lines drawn from the *Planet* to the *Sun*) will also be equal among themselves, or to one another.—It must therefore follow, that when the *Planet* is near its *Perihelion*, its Motion will be so much fwifter than in the Aphelion, as the Line PS is shorter than AS; because the Line PS multiply'd into the Arch PC, must be equal (by this surprising Law) to the Line SA multiply'd into the Arch BA.

The Inclination and Nodes of the Planets Orbits.

Though the Planets revolve at different Distances, one beyond the other, round the Equatorial Parts, or Middle of the Sun, yet they do not all move exactly in the same Plane; but have their Orbs a little inclin'd to one another.—Thus, suppose we make the Plane of the Earth's Orbit the Standard, then we shall find that the Orbits of the rest of the Planets cut it in two opposite Parts, having one Half of their Orb above, the other Half below it. The Points where the Orbits cut the Plane of the Earth's Orbit, are call'd the Nodes of the Planets. No Planet can therefore be in the Ecliptic (i. e. exactly level with the Plane of the Earth's Orb) but when it is in the Nodes: In all other Parts they are either a little above or a little below it. When they are above it, they are said to have North Latitude; when below it, South Latitude; which Latitude is never but a few Degrees. For when the Planets are at the greatest Distance above or below the Plane of the Earth's Orbit, (which always happens when they are in the Middle between the Nodes) the Distance of Saturn is not more than 2½ Degrees; of Jupiter, 1½ Degree; of Mars, 1½ Degree; of Venus, 31 Degrees; and of Mercury, 7 Degrees.—This Inclination seems to have been given the Planets to prevent their too frequently eclipsing one another; which they would have done every Revolution, had not this Divine Contrivance taken Place.

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Lengths of the NATURAL DAY,

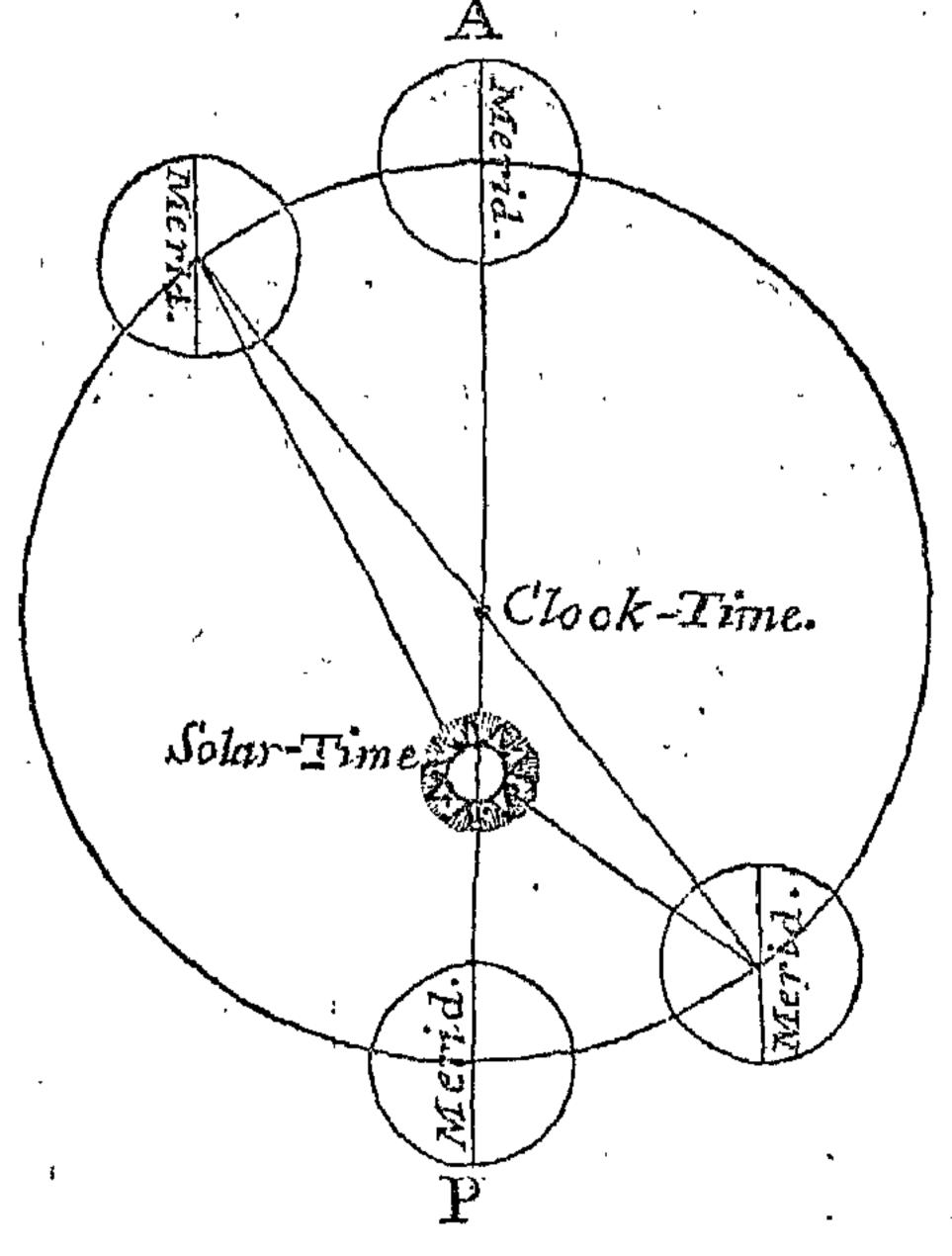
COMMONLY CALLED

THE EQUATION OF TIME.

F the Earth mov'd round the Sun in an Orbit perfectly circular, and her Axis did not incline to her Orbit, every Day throughout the Year would be precise. ly of the same Length: But as her Orbit is a little elliptical, and her Axis inclines thereto, there must be a Difference in their Lengths arising from both these Caules.

With Respect to the former: It is evident from the following Figure,—that as the Earth proceeds from her Aphelion at A towards her Perihelion at P, her Mo-

tion is continually increasing; and she passes over a larger Portion of her Orbit every fucceeding Day than the Day before. Hence it must follow, that any one Meridian will be turn'd towards the Sun's Disk where it is vifibly Noon, sooner than towards the Center of the Ellipsis, which is Noon shown by a Clock; and thereby make the Solar Noon a little precede the Noon, pointed out by the Clock: and this will be the Case thro' this Half of the Orbit.—But in her Motion from the Perihelion at P, to her Aphelion at A, her Motion is continually lessening, and the Arches pass'd over each Day are less and less: Consequently any Meridian will be a little longer in turning towards the Sun, than towards the Center of the Ellipsis; so that thro' this Half of the Orbit,—the Clock-Noon must



always precede the Solar-Noon. - When the Earth is at A and P which happens about the End of June and the End of December, the Clock and Sun are equal; and it is Noon by both at the same Time. The greatest Difference between the Time shown by the Clock and that of a true Sun-Dial (either before or after each other) arising from the Ellipticity of the Earth's Orbit, is never more than about 73 Minutes.

The Difference arising from the other Cause; that is, from the Inclination of her Axis, is owing to the oblique Position of the Equator to the Ecliptic. The

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Revolutions of the Equator, or any of its equal Parts, (on which the Time is always measured by a Clock) are always perform'd in equal Time: But the (Earth or) Sun in passing equal Parts of the Ecliptic, with respect to any one Meridian, takes up sometimes a little more, and sometimes a little less Time. Hence it sollows, that the Solar Days must be unequal among themselves, and differ from those pointed by the Clock. Thro' the first and third Quarters of the Ecliptic; i. e. from Aries to Cancer, and from Libra to Capricorn, the Sun passes the Meridian some standard pointed by the Equator: Consequently the Solar-Noon a little precedes the Noon pointed out by the Clock: But in the second and sourth Quarters of the Ecliptic; i. e. from Cancer to Libra, and from Capricorn to Aries the Sun passes the Meridian a little later than a like Arch of the Equator, therefore the Clock in those Quarters will a little precede the Time pointed out by the Sun.

The greatest Difference arising from this Obliquity of the Equator to the Ecliptic is never more than 9 Minutes 54 Seconds. When the Sun is in the Beginning of Aries and Libra, Cancer and Capricorn, the Time shewn by the Clock and Sun correspond in this Account exactly.

The Time shewn by the Sun, or a good Dial is called apparent Time; and that shewn by a good Clock or Watch is called mean or equal Time. The Days measur'd by the Clock are ever the same Length; that is 24 Hours each; but those measur'd out by the Sun are sometimes a few Minutes more;—sometimes a few Minutes less than 24 Hours; the Difference is called the Equation of Time: And a Table resulting from the above two Causes, calculated to every 5th Day in the 2d Year after Leap Year, is generally set in the outer Case of a Watch; by which you may see at any Time how much the Sun is before the Clock, or the Clock before the Sun on that Day you enquire after.

These Differences in the Lengths of the Days take place in all the other Planets more or less, according as their Orbs are more or less elliptical, and their Axes more or less inclin'd thereto.

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THAT RETAINS THE

EARTH AND PLANETS

INTHEIRORBITS.

ROM many Observations made on the System, it is reasonable to believe that every Particle of Matter in it is endued with the wonderful Property of attracting each other: Hence it is, that loose Bodies upon the Surface of the Earth and Planets are kept from flying off by their Rotations round their Axes; and that Bodies projected into their Atmospheres return back to them again. This Attraction being greater in the Sun, (as he is immensely larger than all the Planets together,) strongly draws them towards himself; and there they would absolutely descend, if that Tendency was not counteracted by a Force impress'd upon each Planet in a Direction contrary thereto. These two Forces, being always equal, and continually operating—one endeavouring to descend toward the Sun, and the other to fly off at right Angles by the projectile Impulse, causes the Planets to revolve in Orbits round him. And as they are plac'd at different Distances, they will revolve in different Periods of Time. The same power retains the Moons, or Satellites in their Orbs and carries them round their Primaries; and thereby adjusts, balances, and supports the whole System. This Attraction in the Sun and Planets, like all Powers diffus'd from Bodies, grows weaker the further it proceeds from them: It decreases exactly as the Squares of their Distances increase; that is, at twice the Distance it is 4 Times weaker, at thrice the Distance of Times weaker; and so on.

Some Philosophers have supposed that these Effects are produced by a Subtle Matter diffused throe the System, pressing towards the Sun and Planets, and acting as all mechanical Causes do, by contact. But when we find that this Force of Attraction is ever in Proportion to the Quantity of Matter in Bodies, and not the Quantity of Surface, acting as strongly upon the internal as external Parts, it seems to surpass the Power of Mechanism, and to be either the immediate Agency of the Deity, or the Effect of some Law originally impressed upon all Matter by Him. And that this Force was at first given by the Deity is evident; for mere Matter

can never put itself in Motion;—and all Bodies, we know, may be mov'd in any Direction whatever; yet the Planets both primary and secondary move the same Way, i. e. towards the East, and nearly in the same Planes, whilst the Comets move in all Directions and Planes very different from one another: Now, these so very different and contrary Motions can never be owing to any Mechanical Cause,—to Chance, or Necessity, but to the Free-Will, Power, Wisdom, and Goodness of an intelligent Being.*

* A very learned and ingenious Author has obliged us with a Computation of the moving Force that is every Instant produced in our System, by the mutual Gravitations of the Bodies that compose it, which may fairly be consider'd as the instantaneous Essicacy (in one Article) of that mighty Cause, whose Power form'd, and whose incessant Activity maintains, this great and beauteous Fabrick. From his Computation, a Force of Motion is every Instant produc'd afresh in the Solar System exceeding 2552604000000000000 Hundred Weight Avoirdupoize. - Enormous Sum!——The Reader perhaps will be surpris'd should I only tell him that if he would attempt to count this Number of Hundred Weights, and would Spend 10 Hours at it every Day, and were to count 100 in every Minute, he could not hope to finish it within the Compass of his Life. But perhaps he will hardly credit me, when I assure him that were the Work transmitted from Father to Son it would employ many Millions of Generations; as at the Rate I have suppos'd (which is Such, at the lowest Estimation, is the instantaneous Production of moving Force,—such has been the Production of every successive Instant ever since the Heavens and the Host of them were finish'd; and such must be the incessant Production of every Instant to come, so long as they shall fublist in their present Form.

Forces of the like Kind must every Moment be produc'd in each of those innumerable Systems, which are only known to the Tenants of this small Globe, by the Glimmerings of their distant Suns. The whole instantaneous Production, therefore, in this System, can be but a small, an infensible Part, of the whole Force that is produc'd every Instant in the boundless Universe.

Sir Isaac Newton, in the Scholium with which he closes his immortal Work, the Mathematical Principles of Natural Philosophy, concludes from what he had discover'd of the Structure of the World, and some of the plainest Metaphysical Axioms,—the Intelligence, the Wisdom, the Eternity, the Ubiquity, the Unity, and the Dominion of the first Cause,—I know Nothing that so sorcibly evinces the Infinitude of his Power as the demonstrable Exertions of it in this single Article. For what can that Power be less than infinite to whose instantaneous Production (in that single Instance the Production of Motion) the vast computable Effect in this System bears but the Proportion of finite to infinitely great?

To give the Learner a better Idea of the various Properties and Affections of the Planets mentioned in the foregoing Course of this Work, I have brought them all into one View in the following TABLE.

Planets Names	Saturn	Jupiter	Mars	Eartu	Venus	Mercury
Characters or Marks	ъ	24	3	0	ç.	¥
Diameters in Miles	67870	81155	4444	79.4	7906	2460
Circumfer. of Bodies	213112	254908	13960	25020	24523	7724
Mean Distance from the Sun, in Miles	777.000.000	424.000.000	1 2.3000,000	81.000.000	59,000,000	32,000,000
Periods, or Length of their Years	10759 ^d 6h 36'	4332 ^d 12 ^h 20'	185° 23° 27'	3654 6h 9	224 ^d 16 ^h 49	87d 23h 16'
Diurnal Motion, or Length of their Days	unknown	od 9h 56'	1d 00h 40'	od 23h 56' Sun 25d 1	od 23h 00'	unknown
Moons, or Satellites	5	4	unknown	I	ſuppos'd 1	not known
Proport. of Light and Heat, supposing the Light at the Earth 1	<u>I</u>	I 27	4	I	2	7
Inclination of their Or-	2° 30′ 00″	1° 20′ 00″,	1° 52′ 00′′	'0° co' oo''	3° 24. 00″	6°.54′ 00″
Inclination of their Axes to their Orbits	30° 00' 00"	little if any	little if any	Earth 23° 29'00' Sun 8 Deg.	little if any	unknown
Daily mean Motion in the Ecliptic	6° 02' 00"	00.04' 59"	0° 31′ 27″	0° 59′ 08′′	1° 36′ 08″	40 05' 32"
Greatest appar. Diam.	0 19" 40"	0 24 12 12 12	0' 20" 50"	32' 47" Sun	1'05' 58"	0'11"48"
Least apparent Diam.	0' 14" 11"	0' 14" 36"	0' 12" 46"	31' 40"Sun	0'09"34"	0' 04" 04"
Eccentricity of Orbits, supposing Dist. 1000	55	48	93	. 17	7	210
Place of Aphelion, the Perihelion opposite	15° 0° 06′	<u>∽</u> 10° 58′	17% I ° 55	VS 8° 57'	70 37	£ 13° 45'
Place of Ascend. Node, Descending opposite	95 21° 26′	्रात 8° 3 2′	8 180 09	* * *	II 14° 34′	8 1,5° 45'
Elongation of Inferior, and Parallax of Supe- rior Planets	6° 00'	110 05	41.00	承 朱 梁	46° 41′	22° 45'
-Colours of the Planets	dim lead-colr.	splendid white	fiery red	* * *	yellowish wh.	fparkl ^g . red
Motion in an Hour, in Miles	18000	240CO.	45000	56000	700co	c00001.
Distances deduc'd from Transit of Venus	907.956.130	.494.990.976	145.014.148	95.173.000	68.891.4.86	36.841.468

The Distance of the Moon from the Earth is about 240,000 Miles.—She revolves round her Orbit, from Sun to Sun, in 29^d 12^h 44'.—Her Diameter is 2175, and her Circumsterence 6831 Miles.—The length of their Day is almost 30 of ours, and their Year the same as with us.—She moves in one Hour 2200 Miles,—Greatest apparent Diameter 33' 20', and least 29' 48".—The Place of the Nodes is always varying, revolving round the Ecliptic in 18' 224^d 4^h.

—Her Eccentricity is 55 Parts in 1000.—The Moons of Jupiter and Saturn are supposed to be as large as our Moon, or some of the Inserior Planets.

Of all the Comets, the Periods, Distances, &c. of only Three of them are known.

Years

Ift Comet revolves { 75 at the 13000 in an 2000 from 2000 Mill. } A8000000 from 2000 Mill. }

Ift Comet from 2000 Mill. }

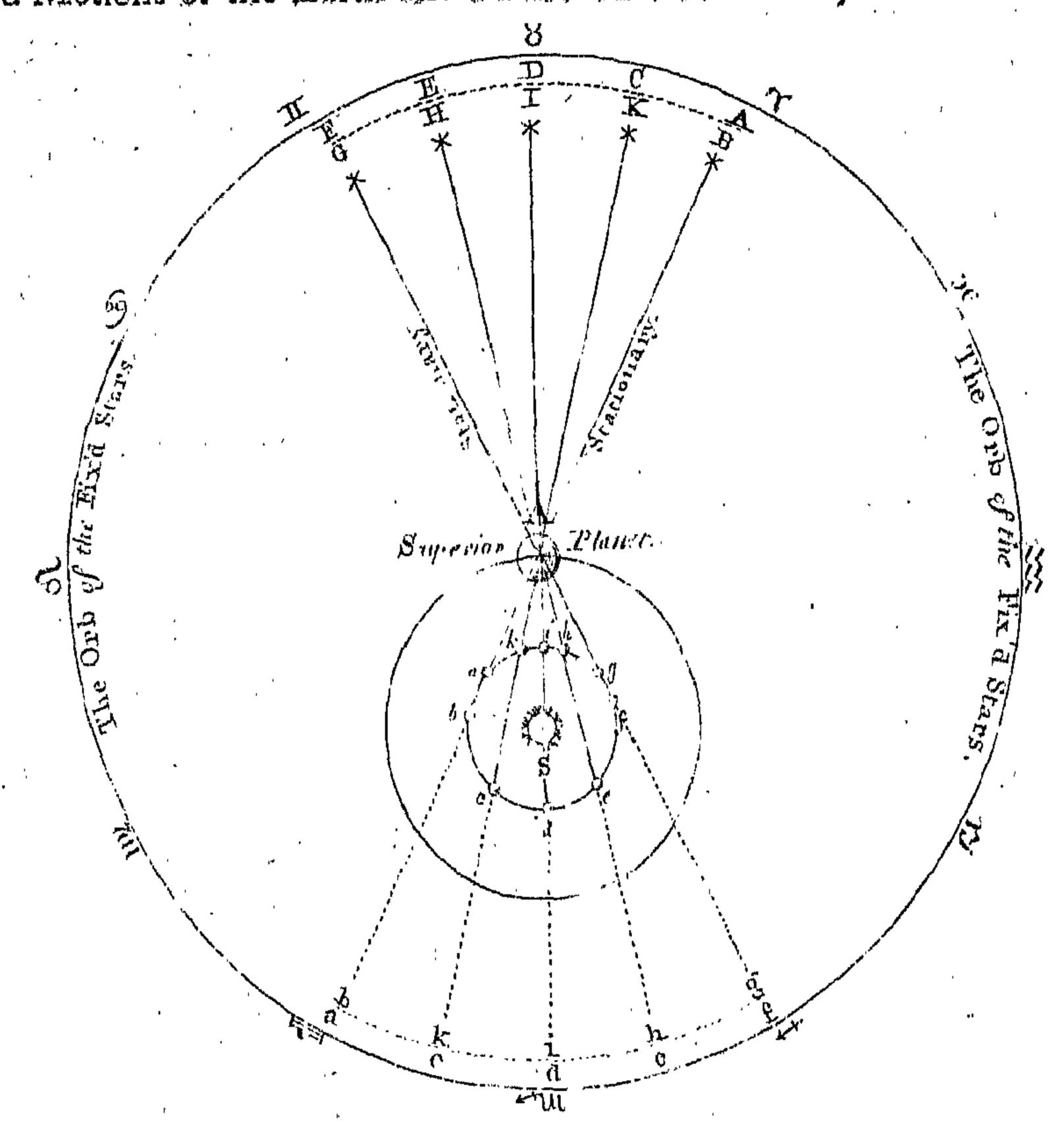
Some Comets are as small as the Moon, and others as large as our Earth.

The Distance of the nearest six'd Star is about 20.000.000.000 Miles; and they seem to make one Revolution on the Poles of the Ecliptic in 25920 Years.

OF THE

STATIONS AND RETROGRADATIONS. Of the PLANETS.

If the Planets were viewed from the Sun, they would always appear to move round him direct, and according to the Order of the Signs, i. e. from Aries to Taurus, &c. but to us, who see them from the Earth, they appear sometimes to move on direct,—sometimes to be retrograde, or to run back,—and sometimes to be stationary, or fixed to one particular Point in the Heavens. This Diversity of their Appearances is owing to the united Motions of the Earth and Planets in their Orbits, as here delineated.



In this Scheme, let S represent the Sun, the Center of the System; the smaller Circle a, b, c, d, &c. the Orbit of the Earth;—and the larger the Orbit of one of the superior Planets;—also let the outer Circle v, v, v, ecc. represent the Orb or Sphere of the fix'd Stars, where the Place and Motions of the Planets are determined. Then it is evident, that tho' the Planets all revolve round the Sun the same Way, and with a regular Velocity, yet to us on the Earth (and to one another) their Motions will seem very different. For sometimes they will appear to move slowly, sometimes to stand still, and at other Times to run back, or move contrary to their real Motions. Thus, suppose the Earth in her Orbit at a ;—then an Eye will behold the superior Planet A (suppose it be Saturn) in the Line continu'd from the Earth, by the Planet, to the Stars at A. All the Time the Earth passes from a to b, the Planet will appear to stand still, or six'd in the same Part of the Heavens, because the Line proceeding from the Earth to the Planet continues the same.—As the Earth moves from b to \bar{c} , the Planet will appear to move from B to C, with a direct Motion among the Stars, also.—At D, he is said to be in Conjunction with the Sun, i. e. in the same Place of the Heavens with that Luminary; and, if in the Plane of the Ecliptic, will pass behind his Disk, and rise and set when he does.—As the Earth goes on to e and f, the Planet will be seen to move direct to E and F likewse: But whilst the Earth passes from f to g, the Planet will appear to be flationary, or not to move at all.—And as the Earth revolves to b, the Planet will seem to run back to II.—At i, he will appear at 1; where he is said to be in Opposition to the Sun; he now rises as the Sun sets, and

Thines

thines all the Night.—As the Earth proceeds to k, the Planet continues to go retrograde; and when the Earth is come to a, the Planet becomes stationary, as at first.

From this Scheme it is also evident that the same Appearances must happen to the inferior Planets as to the Superior. For an Inhabitant at the Superior Planet at A, will see the inferior move on direct,—become stationary and retrograde, at the same Time, and in the same Manner, but in the opposite Part of the Heavens.— Thus, whilst the superior Planet is beheld from the inferior, at a and b, to be stationary at AB in v, the inferior will be seen from thence stationary at ab in -.-When the inferior is at c, he will appear in the opposite Part at c.—At d he will appear at d, conjoin'd with, and in the same Part of the Heavens with the Sun. This Situation of an inferior Planet, because it is situate beyond the Sun, is call'd its superior Conjunction; and if it has no Latitude at that Time, will pass behind the Sun's Body, and suffer an Ecliple.—As he advances from d to e, he appears, from A, to move on still direct from d to e.—At f, he is seen at f, where is his greatest Elongation on that Side his Orbit from the Sun, and will set in the Evening after him.—Whilst the inferior moves from f to g, he appears stationary, and (as it were) fix'd in the same Part of the Heavens, just as the Superior does from thence in the opposite Parts; because the Lines, passing from each to each, terminate in the same Places during that Time.—But, as he proceeds from g to h, he appears to the superior (as the superior does to him) to go back or retrograde in the Heavens. —At i, he will be found at i, conjoin'd with the Sun again. This is call'd his inferior Conjunction, because hetween the superior and the Sun; and, if he is exactly in the Plane of the Ecliptic, will pass over the Sun's Disk, and appear like a black Spot on his Face.—As he revolves from i to k, and a, he fill recedes backwards, and appears amongst the Stars at k and a, from a to b he appears stationary, as well as the superior, and for the same Reasons as they did, on the other Side of the Orbit, at f and g.—At a he is said to be at his greatest Elongation on this Side the Sun, and sets before him.—At b he begins to move direct as before.

I have considered the superior Planet as remaining in the same part of its Orbit during one Revolution of the Earth, in order to convey the Ideas of their Stations, Directions, and Retrogradations easier to the Learner. For the they actually advance, at the same Time, the same Way, but as the Motion of the superior is much slower, it may, in this Illustration, be safely omitted. For instance, while the Earth revolves once round,—Saturn will have gone but one thirtieth Part of his Orbit; Jupiter in the same Time, would have gone about one twelfth Part; and Mars about a Half.—Whence it follows, that these Appearances happen oftener in Saturn than in Jupiter; and oftener in Jupiter than in Mars; because the Earth sooner overtakes (or comes up with) the former of these Planets than the latter, whose Motion is faster. And, for the same Reason, they happen oftener in Mercury than in Venus, because he circulates not only faster, but in a less Orbit, and therefore oftener overtakes the Earth. These Retrogradations are greater, the nearer the Planets are to the Earth.

other;—if, on the contrary Side, they always appear direct;—but for some little Time before and after their Conjunctions, they appear stationary. This is an undeniable Proof of the Earth's Motion, and that the present System is the true System of the World: Because if the Earth be supposed to remain fixed in the Center of the Universe, these various Motions could never take place here; but all the Planets would appear to move (as they really do) the same Way, and revolve (like the Moon) regularly round us.

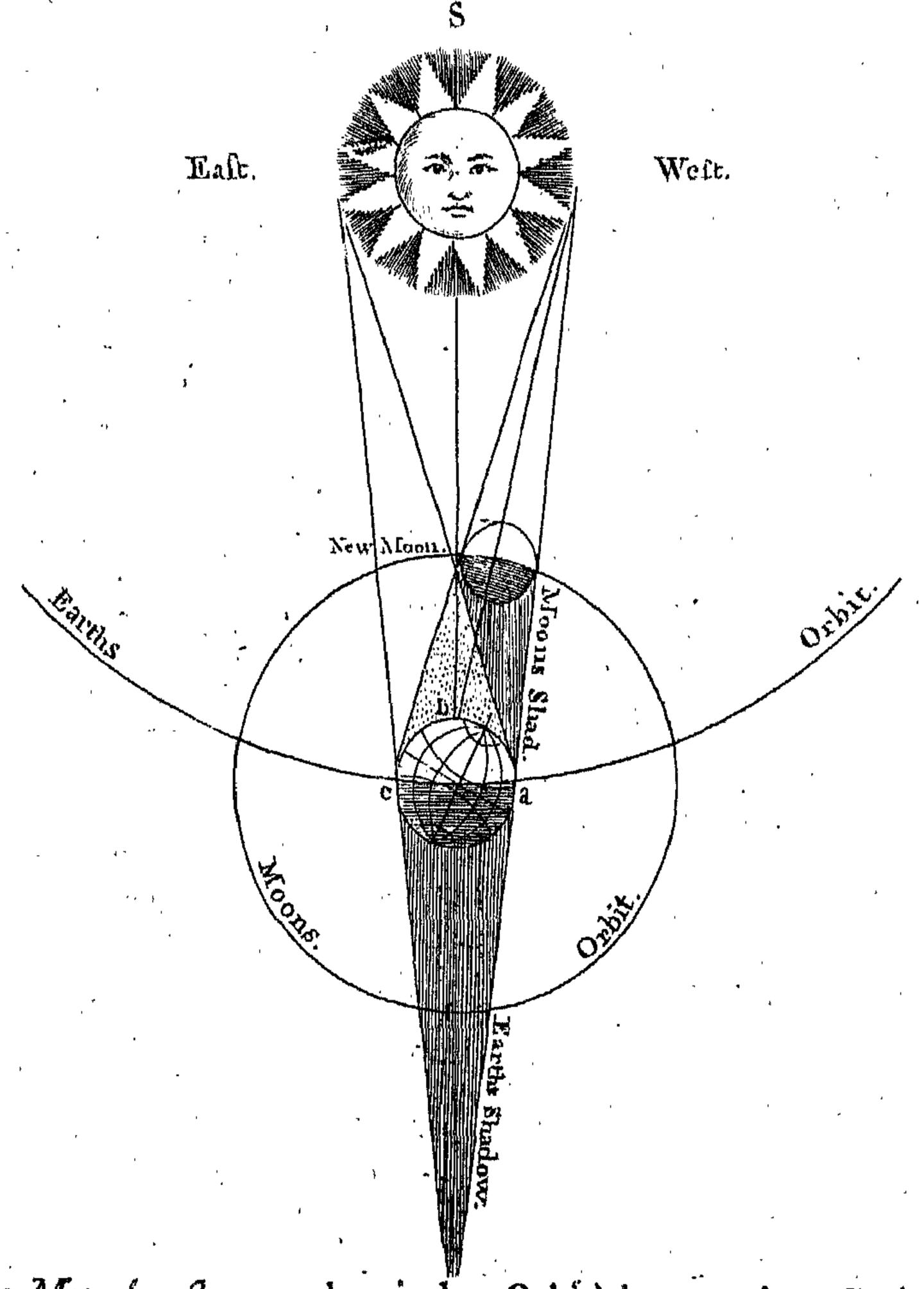
OF

Of ECLIPSES.

N Eclipse is that Obscuration, or Deprivation of Light, in any of the heavenly Bodies, which is caused by some other coming between the Sun and that Body. There are various Kinds of Eclipses in the Heavens, but those of the Sun and Moon are most remarkable.

Eclipse of the Sun.

HE Eclipse of the Sun (or more truly of the Earth) is caus'd by the Moon, coming between the Sun and Earth, and by that means hiding his Light from us. This can never happen but at the New-Moon, as is evident from an Inspection of the following Scheme; in which S represents the Sun, E the Earth, and M the Moon in her Orbit revolving round the Earth.



Now when the Moon (as she revolves in her Orbit) is come into Conjunction with the Sun, i. e. just between Us and Him, which can never happen but at the New-Moon, it is manifest, that her Shadow will fall upon the Earth at a, and by that means hide his Face and Light from those that live upon that Spot. But as the Moon is much less than the Earth, the Shadow of the Moon cannot cover the whole but only a Part of the Surface next it. For this Reason, you see, that the same Eclipse of the Sun may be total to one Country;—to another partial or Part eclips'd:—and to a third, none at all.

Thus in the foregoing Scheme, it is manifest that to the Inhabitants of that Part of the Earth marked a, the Sun will seem to be totally dark, they being wholly involv'd in the Shadow.—To those at b, he will appear to be eclips'd in one Part only, and the more so—the nearer the Spectator is to a, they being immers'd only in the partial Shadow—And to those that live at c, he will not be eclips'd at all*.—At the same Time an Inhabitant at the upper Part of the Moon, (if there be any) will see the Sun sree from any Obscurity, whilst those on the under Side will behold the Earth, Part of it involv'd in a dusky Shade.

The Sun's Eclipse always begins on the Western and ends on the Eastern Side his Body. The Reason is,—because the Moon (which is always the Cause of this Obscurity,) moving round her Orbit from West to East, must necessarily first arrive at, and touch the Sun's Western Limb, and go off at the Eastern.

As the apparent Diameters of the Sun and Moon are nearly equal, the total Duration of an Eclipse of the Sun can never continue more than two or three Minutes; but the same solar Eclipse, from Beginning to End, may be 2. Hours.—

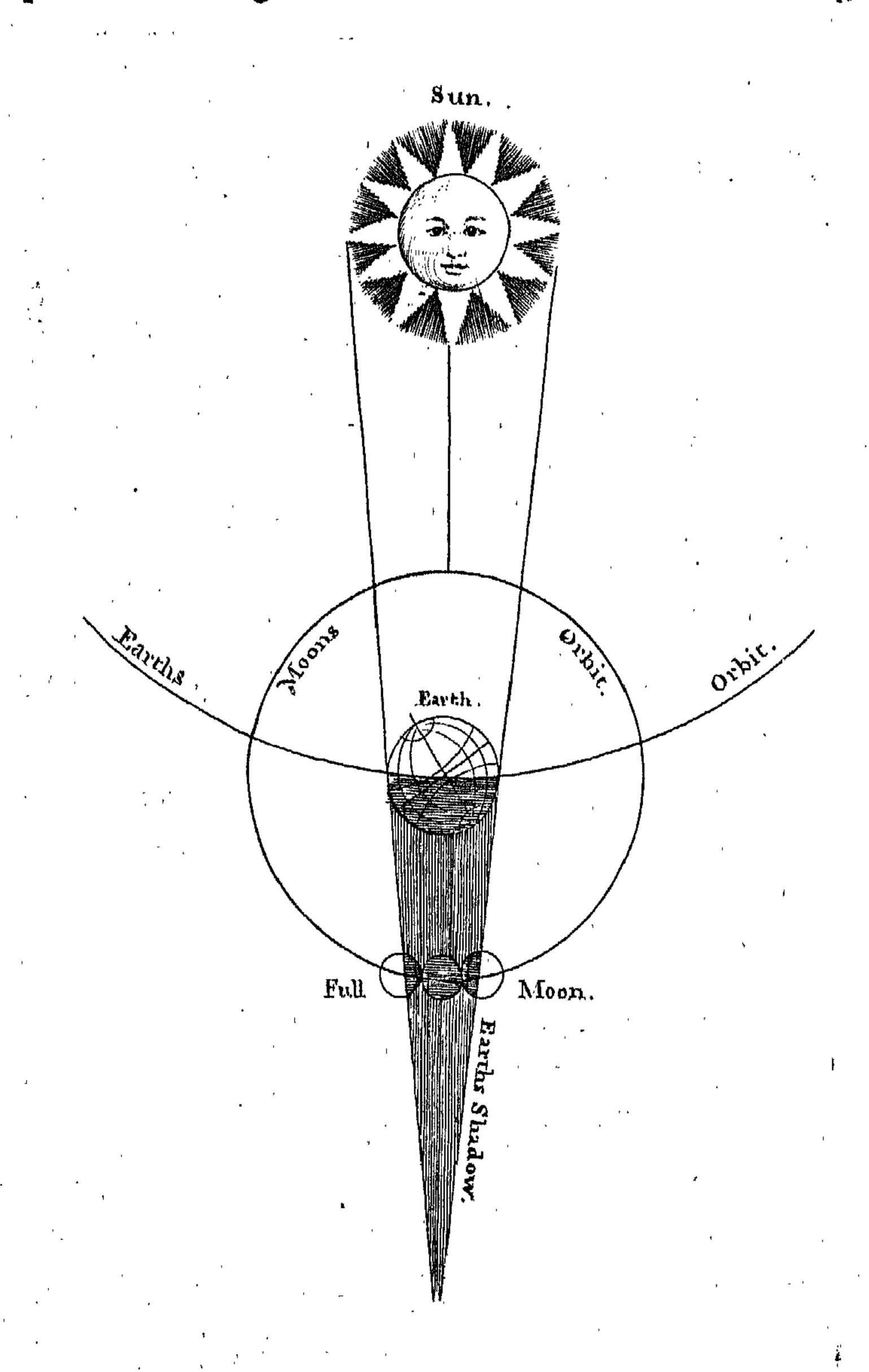
In order to determine the Quantity of an Eclipse, Astronomers have divided the Sun's Diameter in 12 equal Parts, call'd Digits; and therefore, when they say the Sun will be eclips'd 5 or 6 Digits, they only mean, 5 or 6 Fingers Breadths of the Sun will be obscur'd or darken'd, whilst the rest will continue bright and visible to the Spectator.

An Eclipse of the Sun can only be visible by Day, when the Sun is above the Horizon: and though there be 12 New-Moons every Year, yet it is seldom we have more than 4 Eclipses of the Sun, the Reason of which will be shewn 2 or 3 Pages further on.

- * The Shadow of the Moon will cover a greater or less Portion of the Earth's Surface, as the Moon happens to be nearer or further from us at the Time of an Eclipse. At a mean, the Shadow may spread itself over 180 Miles in Diameter, and moves at the Rate of 2100 Miles in an Hour. But the partial or penumbral Shadow, which is that faint Shade surrounding the Shadow of the Earth, will extend the Breadth of 4900 Miles, and be five Hours in passing over the Earth's Disk. As soon as that Shade touches the Earth the Eclipse begins; and when it wholly leaves it the Eclipse ends. The Moon's Shadow travels over the Earth towards the East at the Rate of 36 Miles a Minute, which is about 4 Times swifter than a Cannon Ball, when first discharg'd from the Mouth of that Piece.
 - In the early Ages of the World, the Nature and Cause of Eclipses was so little known, that the Ignorance and Superstition of those Times looked on them as Prodigies, and portentous of some great and malevolent Effects: But the Learning of the present Age has taught us, that these Appearances happen according to the Course of Nature, and are the necessary Result of those Laws which the Divine Being gave them at the Creation.—Ricciolus, in his Almagist, relates a Story of Columbus, who being drove on Shore at the Island of Jamaica, in 1493, and meeting with great Distress and Cruelty from the barbarous Inhabitants, threaten'd to afflict them with some grievous Calamity; assuring them as a Mark of his Power, that on such a Day the Sun should be darken'd; which by an Eclipse, happen'd on the Day he had assign'd. This surprising Phænomenon wrought such Terror in their Minds, that they looked on him as something more than human; and, therefore, they afterwards did all in their Power to serve him.

ECLIPSE of the Moon.

N Eclipse of the Moon is caus'd by the Interposition of the Earth between the Sun and Moon: This can only happen at the Full-Moon, when she is directly opposite the Sun; for the Earth, being then exactly between the Sun and the Moon, will cast her Shadow upon the Moon, just as the Moon did her Shadow upon the Earth before, in the Eclipse of the Sun. This is clear from a Sight of the solution figure; where it is evident, that when the Moon is arriv'd to the opposite Part of her Orbit, she falling, more or less, into the Earth's Shadow, is by that Means depriv'd of her Light from the Sun, and so suffers an Eclipse.



If only Part of the Moon's Body passes through the Earth's Shadow, she is said to suffer a partial Eclipse.—If the whole Body is immers'd in the Shadow, and begins immediately to emerge again, the Eclipse is said to be total without Continuance.—But if the Moon passes through the Middle of the Shadow, (or near it) her Stay there will be considerable; as the Shadow, in that Part of it, is about three Times broader than the Moon, and the Eclipse will be total with Continuance.

An Eclipse of the Moon appears to all Parts where it is visible to be the same in Quantity and Duration as it really is. For the Moon being an opaque Body, shining only with a reflected Light from the Sun, is deprived of that Light by being immers'd in the Earth's Shadow*.

The Moon's Eclipse begins always on the Eastern Side, and ends on the Western, contrary to that of the Sun. For her Eclipse being real, and by Reason of her direct Motion always Eastward, the Eastern Part of her Body must necessarily first touch the Earth's Shadow, and the Western Part leave it last.

The Shadow of the Earth, where the Moon passes through it, being almost three Times as large as the Diameter of the Moon, it will often happen that total Eclipfes of the Moon may be of very different Lengths; but it is seldom that total Darkness continues more than 1\frac{1}{4} Hour, or the whole Eclipse, from Beginning to End, more than 3\frac{1}{4} Hours.

The Body or Disk of the Moon is supposed to be divided (like that of the Sun) into twelve equal Parts, or Digits; and as many as are the Parts immers'd in the Earth's Shadow, so many Digits is the Moon said to be eclips'd. But as the Diameter of the Shadow is far greater than the Diameter of the Moon, it must follow, that, when the Moon is immers'd in the Middle of the Shadow, not only the twelve Digits are eclip'd, but there remains a considerable Part of the Shadow to be pass'd thro' still. In Lunar Eclipses, therefore, which are total, we reckon the Number of Digits according to the Quantity of the Earth's Shadow over, or beyond the Moon, when it is immers'd wholly in the Shadow, which may sometimes extend to 23 Digits nearly.—All above 12 shew how many Parts of the Shadow the Moon has to pass thro' before she begins to appear on the opposite Side.

The Moon's Eclipse is visible only by Night, i. e. when she is above the Horizon: And though we have twelve Full Moons within the Circle of the Year, yet never more than three of them suffer an Eclipse. The Reason you will find in the next Pages.

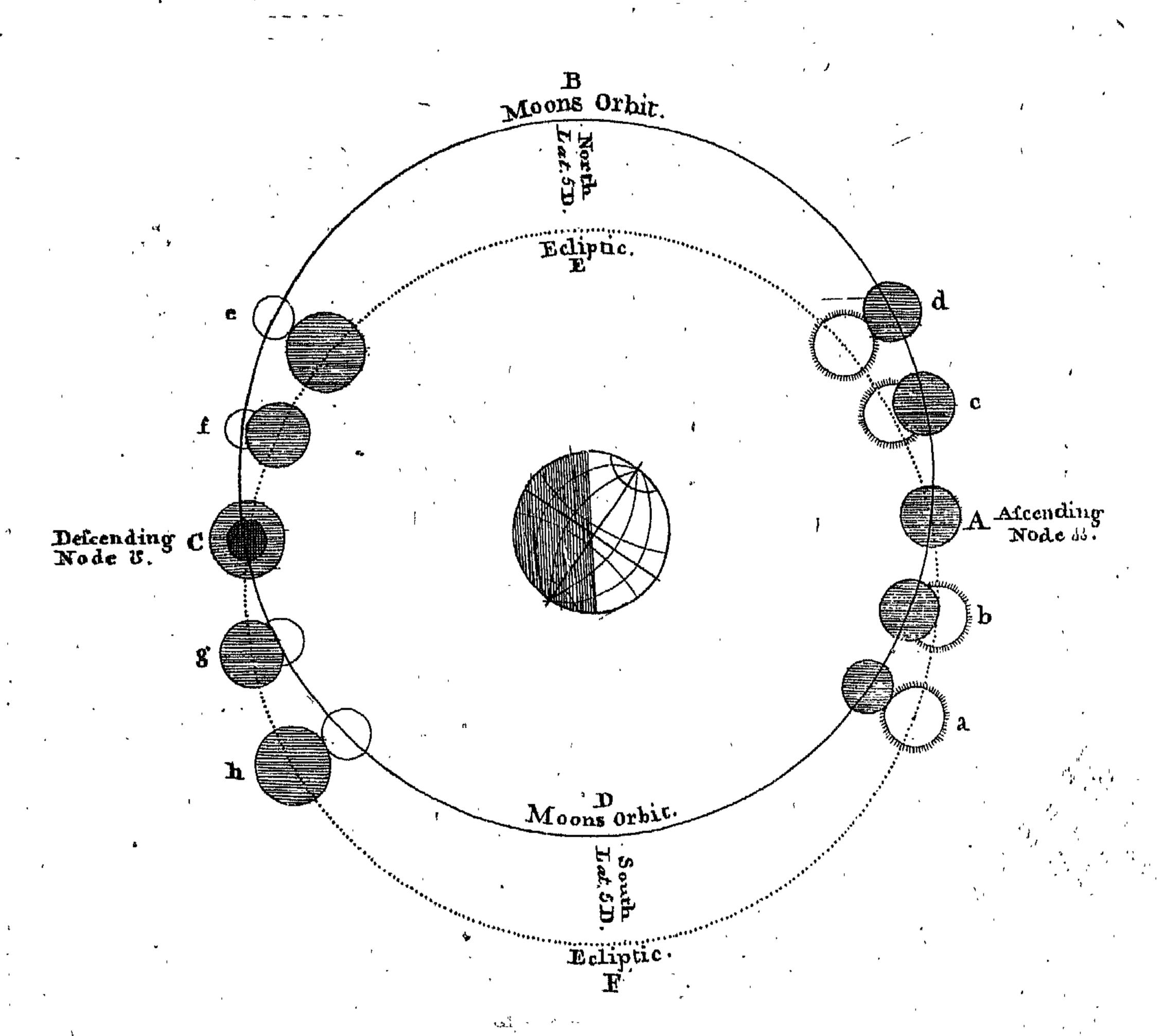
* All the Inhabitants of the Moon on the Side next us, that fall into the dense Shadow of the Earth, have a total Eclipse of the Sun; but all those within the Penumbral Shade, see only a partial Eclipse of his Body. In total Aclipses of the Moon, when she is in or near the Center of the Earth's Shadow, her Body is frequently seen of a languid swarthy Colour: This is owing not to any Light she has in herself, but to the Rays of the Sun falling into our Atmosphere, which being there-by refracted, are bent out of their direct Course into the Shadow: These Rays falling upon the Moon, when she is even in the Middle of the Shadow, render her visible to the Spectators at that Time.

rom the Writings of the Apostles we learn, That at our Saviour's Crucifixion there was Darkness over all the Land from the sixth to the ninth Hour, i. e. from 12 to 3 in the Asternoon; which Darkness some Persons may, perhaps, attribute to an Eclipse of the Sun. But that is impossible: For Christ suffer'd at the Time of the Jewish Passover, which was always celebrated at the Full-moon, when these two Luminaries are in Opposition to, and at their greatest Distance from each other. Besides, such uncommon Darkness could not happen from any natural Eclipse of the Sun: For as the apparent Diameters of these two Planets are nearly equal, such Eclipse could not last but a sew Minutes, whereas this Darkness lasted about three Hours.

Dionysius, a Judge of Areopagus, near Athens, being at Heliopolis, and observing this practernatural Phanomenon, cry'd out, that Nature was dissolving, or the God of Nature suffers. He afterwards embrac'd the Christian Faith (from this Astronomical Conviction,) and suffer'd Matyrdom for the Truth of it.

Of the Limits of Eclipses both Solar and Lunar, with the yearly Number of each, and the Times when they happen.

HE Reason why the Sun is not eclips'd at every New-Moon, nor the Moon at every Full, is because the Moon's Orbit does not lie in the same Plane or Tract with that of the Earth's, which is call'd the Ecliptic Line, but cuts or crosses it in two opposite Points, making an Angle therewith of about 5 Degrees. These two Points or Intersections are call'd the Nodes of the Moon. The Point where the Moon begins to ascend above, or become Northward of the Ecliptic Line, is call'd the ascending Node, and is usually mark'd by Astronomers thus (S). The other Point, from which the Moon begins to descend below, or go Southward of the Ecliptic, is call'd the descending Node, and is mark'd thus (S). See the following Scheme.



In this Scheme let the larger black Circle, ABCD, represent the Moon's Orbit with the Earth in the Center: And let the dotted one, AECF, represent the Ecliptic, or Plane (at the Distance of the Moon) in which the Sun and Earth's Shadow appear to move. Then will A and C be the two Nodes, or Knots where the Moon's Orbit and Ecliptic intersect each other in an Angle (at A and C) of 5 Degrees; one half of the Moon's Orbit, ABC, being above the Ecliptic, the other half, CDA, below it. The little Suns upon the dotted Circle, about the Node A, represent the Sun in those Parts of the Ecliptic; and the little shaded Circles are the New-Moons, at those Places also.——On the opposite Side, the

black Circles represent Sections of the Earth's Shadow, which are always opposite the Sun; and the lighter Circles are the Moons, which happen there at the Full.

Now suppose the Moon, in passing round her Orbit (which she does every Month), comes into Conjunction with the Sun at a, where she is below, or Southward of the Ecliptic Line, then it is apparent (from the Figure) that her upper Limb will just touch the Sun's lower Limb, without eclipsing it.—But if the New Moon happens at b, then the Moon will pass over the lower Part of the Sun's Disk, and cause an Eclipse.—If at the Time of the New-Moon, the Moon meets with the Sun in the very Node, at A; then there will happen a central and total Eclipse of the Sun. But as their apparent Disks are nearly equal, the total Darkness will continue but two or three Minutes at most.—If the New-moon happens at c, then you see there will be a partial Eclipse of the Sun; and the Sun's upper Part will be obscur'd, as the Moon is now got Northward, or above the Ecliptic Line.——If the Luminaries meet at d, then the Moon's lower Limb will touch the Sun's upper Limb, but pass by without eclipsing him.

Again; As the Moon passes from the Sun to the opposite Part of her Orbit, where she is in the Full, as at e, then her lower Limb will just touch the Edge of the Earth's Shadow, but pass over it without being, in any Part, eclipsed.—But if the Full-moon happens nearer the Node, at f, then the lower Part of the Moon dips into the Shadow, and causes a partial Eclipse.—But if the Full Moon falls at the very Node C, the Moon passing there through the Middle of the Shadow will be totally and centrally eclipsed; which Eclipse will be of some Continuance, as the Shadow is, in that Part, nearly three Times the Diameter of the Moon's Body.——If the Opposition, or Full-moon happens at g, then there will be a partial Eclipse; but the upper Part of the Moon will be immers'd in the lower Part of the Shadow, because she is now below the Ecliptic.——At b, the Moon's full Disk passes under the Shadow free from any Obscuration.

The Distance of the Moon, in Degrees and Minutes, above or below the Ecliptic Line, is call'd her Latitude. If she is above the Ecliptic, she is said to have North Latitude; if below it she has South Latitude: And if the Latitude, at any Time, exceed the Sum of the Semidiameter of the Moon, = $10\frac{1}{4}$ Minutes, and the Earth's Shadow = $45\frac{2}{4}$ Minutes, the Moon, at that Time, cannot be eclips'd; but will either pass under or over the Shadow, just as she happens to be above or below the Ecliptic Line.—The Distance from the Node, either before or after it, corresponding to the above Latitude, is about 12 Degrees; consequently, that is the Limit of Lunar Eclipses For when a Full-moon happens within 12 Degrees of the Node, then she will be eclips'd; and the nearer the Node, the greater will the Eclipse be.

If at the New-moon, the Latitude of the Moon exceeds the Sum of the Semi-diameters of the Sun, = $16\frac{1}{4}$ Minutes, and the Moon, $-16\frac{3}{4}$ Minutes, we should see no Eclipse of the Sun, if beheld from the Center of the Earth:

But as we view the Luminaries from the Surface, which is much higher, we are obliged to take in the Semidiameter of the Earth, as feen from the Moon likewise, — 61½ Minutes. Then, I say, if the Latitude of the Moon be greater than the Sum of these three Numbers, — 94½ Minutes, the Sun will not be eclips'd; for the Moon will pass either over or under his Disk, according as she is above or below the Ecliptic Line. ——The Distance from the Node, on either Side, agreeing to this Latitude of the Moon, is about 18 Degrees which is the utmost Limit of solar Eclipses: Whence it follows, that if the Sun and Moon, at the Time of New-moon, happen to be within 18 Degrees of the Node, the Sun will suffer an Eclipse at that Time, and the greater—the nearer they are to the Node.

If the Sun should happen to be eclips'd on one Side the Node, just within the Limit between a and b; then the following Full-moon will be eclips'd, but not in the opposite Degree of the Ecliptic exactly,—but something nearer the Node C; because the Earth's Shadow is advanc'd about 14 Degrees forwarder in the Ecliptic in that Time. And the returning New moon will overtake the Sun, and ecliple him again, somewhere between c and d.—But if an Eclipse of the Sun happen at, or near the Node A, then the following Full-moon will not arrive at the Earth's Shadow (which is still moving forward) till it is beyond the ecliptic Boundary at b; so that there can be no Eclipse at that Full.—The other New and Full-moons will pass without being productive of any Eclipse, till the Sun arrives to the oppofite Node C; where the same Number and Kind of Eclipses, which fell at the former Node, will fall again at this. For the Sun and Earth's Shadow having only chang'd Places, the same Eclipses must be repeated, in the same Manner, and for the same Reasons, as before. Hence it is manifest, that there can never happen more than six Eclipses in any Year, i. e. two of the Sun and one of the Moon at each Node; nor less than two, i. e. one of the Sun at each Node, without any Lunar one.

As Eclipses only happen in those Months in which the Sun enters the Signs where the Moon's Nodes are situate; and as the Nodes are directly opposite, the Times of Eclipses will happen at opposite Parts of the Year. Thus, if any Eclipse happens in January or February, then will also happen another in June or July; the Sun, in those Months, being in Signs of the Ecliptic opposite to one another.

Of the PERIOD of ECLIPSES,

Both Solar and Lunar.

F the Places of the Moon's Nodes were fix'd, Eclipses would always happen nearly at the same Seasons of the Year: But as they have a Motion of about 3 Minutes 11 Seconds every Day backward in the Ecliptic, or contrary to the Order of the Signs, i. e. from Aries to Pisces, &c. the succeeding Eclipses must recede backward, from March to February, likewise: And in one Revolution of the Nodes, which is completed in 18 Years, 224 Days, 3 Hours, they will revolve in a retrograde Manner thro' the Year, and return to the same Places again.

But there is a more correct Period of Eclipses discover'd by Observation*; which is 18 Years, 11 Days, 7 Hours, and 43 Minutes: For in that Time, the Sun and Moon advance just as far beyond a compleat direct Revolution in the Ecliptic, as the Nodes want of compleating their retrograde one. Consequently, as the Sun and Moon meet the Nodes at the End of that Period, the same solar and lunar Aspects, which happen'd 18 Years, 11 Days, 7 Hours, and 43 Minutes ago, will return, and produce Eclipses of both Luminaries, for many Ages the same as before.

—On this Principle the following Table and Calculations are constructed.

- The same Kind of Eclipses happen to Jupiter and Saturn: But as Jupiter has four Moons, and Saturn five, and as their Revolutions are much quicker than ours, Eclipses must happen more frequently there than here.
- *** The Eclipses of Saturn's and Jupiter's Moons will be consider'd and treated of in another work.

The USE that may be made of ECLIPSES is very great, not only to the Astronomer and Chronologist in ascertaining more accurately the Periods of the Planets, and fixing the ancient Accounts of Time, but to the Geographer and Marriner in determining the Longitude of Places at Sea or Land. For having an Eclipse truly calculated to any particular Meridian, as suppose London, where the Eclipse is found by Calculation to begin at 5^h 32^m, in the Asternoon; and being at Sea, the same Eclipse is observed to begin at 6^h 16^m, the Difference, which is 44^m, turn'd into Time, allowing for every 4 Minutes 1 Degree, makes 11 Degrees, and so much is the Observer to the Eastward of London.—
But if the Time had been found to be sooner than by the Calculation, then is the Person so much to the Westward of London. By this Discovery the Mariner is enabled to pass with greater safety over the Surface of the pathless Deep, for by knowing where he is, he knows the better how to evade the Dangers, and to direct his Course to the wish'd for Port.

^{*} This Period was found out by the ancient Ghaldeans, and by them call'd Saros by us the Chaldean Saros: But it has been found that these Eclipses fall back from the Nodes about 18 Minutes in every Period, they must therefore in Time wear out and fresh ones come in.

ATABLE OF ECLIPSES.

So adapted, that the young Astronomer may, with very little Trouble, calculate the Number of Eclipses in any Year; discover whether they are of the Sun or Moon; with the Quantity eclips'd; and the exact Time of the Day or Night they fall on.

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The Use of the foregoing Table.

DD 18 Years, 10 or 11 Days, 7 Hours, and 43 Minutes to any Eclipse in the Table, and the Sum will be the Return of the same Eclipse.—If there be but 4 Leap Years in the 18 add 11 Days; if 5 Leap Years add only 10 Days.

In these Calculations, the Day is always supposed to begin at Noon, and end after 24 Hours, on the next day at Noon—When the Day is in December, and the Addition of 10 or 11 Days makes it more than 31, then the Overplus must be reckon'd as so many Days in January in the following Year.—The same must be done with Respect to the Months.

Remember to add 11 Days more, for the Alteration of the Style, after September 14th, 1752.

An Example or two will make all plain.—By this Table it appears that the Sun was eclips'd 1745 Sept. 14 5 20 Old Style. Add, because there are 4 Leap Years in the TI 7.43 1763 Sept. 25 13 Gives first Return of that Eclipse Add in Days for Alteration of the Style Sept. 36 13 3 Deduct for the Month September 1763 Oct. 6 13 3 New Style. Gives Add, because there are 5 Leap Years in the next The next Return of the same Eclipse 1781 Oct. 16 20 46 New Style. Hence it is evident, that the Sun will be eclips'd October 16d 20h 46m; that is, the 17th Day at 3 Quarters past 8 o'Clock in the Morning. From the same Table you may calculate backwards, or for Years past.—For by only subtracting 18 Years, 10 or 11 Days, 7 Hours, and 43 Minutes from any Eclipse in it, you will have the Time when that Eclipse happen'd before. Thus the Sun was eclips'd as above 1745 Sept. 14 5 20 Old Style: Subtract, because there were 5 Leap Years in the 18

Happen'd before in Subtract, because only 4 Leap Years in the Happen'd before Happen'd before 1709 Aug. 23 13 54 Subtract, because 5 Leap Years in the 18 10 7 43

Happen'd before 1691 Aug. 13 6 11 OldStyle.

All these were great Eclipses of the Sun, somewhere upon the Earth, as the Moon was (you see in the Table) very near the Southern Node*.

^{*} If the Moon's Latitude be less than 58 Min. the Eclipse will be total to some Part of the Earth: If more than 58, and less than 90, the Eclipse will be partial only: If more than 90 or 92, the Sun will not be eclips'd at all——To find its Beginning and End. From 8100 subtract the Square of the Moon's Latitude; the Square Root of that Remainder, divided by 31, will give Half the Continuance of the Eclipse, in Hours and Minutes; which added to, and subtracted from the Middle, found by this Table, will give you the Time it begins and ends, i. e. when the Shadow first touches the Earth, and when it last leaves it. This Calculation does not determine the Beginning and End to any particular Place, as London, &c. That requires another Method, and is difficult on Account of the Moon's Parallax.—The Shadow of the Moon travels, in solar Eclipses, over the Earth, at the Rate of about 36 Miles in a Minute.

In like Manner you may calculate an Eclipse of the Moon.

By this Table you see the Moon was eclips'd Add, because there are only 4 Leap Years in the	D. H. M. 1744 Apr. 15 8 32 Old Style. 18 11 7 43
Return of that Eclipse Add 11 Days for the Alteration of the Style	1762 Apr. 26 16 15
	1762 May 7 16 15 New Style.

Hence we find that the Moon was eclips'd on May 8, about a Quarter past 4 o'Clock in the Morning; and that her upper Limb was immers'd in the lower Part of the Earth's Shadow, because she had about 35 Minutes, of South Latitude Assending.—If the Latitude of the Moon be less than 26 Minutes the Eclipse will be total: If more than 26, and less than 58 or 60, the Eclipse will be partial only: But if more than 60, there will happen no Eclipse at all*.

By this Period you may find all the Aspects of the Moon throughout the whole Year: For if the 10 or 11 Days, 7 Hours, and 43 Minutes be added to the Fulls Changes, and Quarters of the Moon as found in an Almanack 18 Years before, you will have their Returns on the Days and Months of the present Year.

These are the Laws which GOD, the great Former of the Universe, gave to the Heavenly Bodies at the Creation, which Laws are constant, uniform, and regular without Variation. To arrive to a perfect Knowledge of them is to arrive at the Apex, or Top of all human Learning.— What lofty—what exalted Ideas must we have of the Deity! his Wisdom, Power, and Goodness! when we view the supendous Canopy of Heaven, and raise our Heads in Contemplation to that almost infinite Extension of the starry Firmament;—And when we are there, if we have liv'd good and pious Lives on this our probationary Planet below, our Souls want but a Step or two to penetrate into the Abodes of the Blessed, and be happy for ever.

* To find the Quantity, or Digits eclips'd.——From 58 subtract the Moon's Latitude, multiply the Remainder by 6, and divide the Product by 16, so will the Quotient give the Digits sought.

To find (nearly) the Beginning and End.—————————From 3364 subtract the Square of the Moon's Latitude, the Square Root of that Remainder, divided by 31, will give the Hours and Minutes of Half the Continuance, which subtracted from, and added to the Middle, (found as above) you will have the Beginning and End of that Eclipse.

The Reason of this Rule may be seen at Page 36 of my Trigonometry.

THE

DESCRIPTION AND USE

OF THE

ASTRONOMICAL CLOCK.

The fix'd Part is a Circle divided into the 24 Hours of the Day. The upper XII stands for 12 at Noon, and the lower XII for 12 at Midnight.—The other Part is a moveable Circle, with such Stars and Constellations delineated in it as are near the North Pole, and never set in this Part of the World. The Stars are bounded with the 12 Months, the Days of which are so adapted, as to stand against the Star, which comes to the South, or upper Part of the Meridian, with the Sun, at 12 o'Clock on those Days.



The Use of this Instrument is pleasant and easy. For, only hold up the Book, turning the Back of it towards the North Pole; then move the inner Circle about, till the Stars inserted therein lie in the same Situation with those you see in the Heavens; (as the Great Bear, commonly call'd Charles's Wain, the Little Bear, &c.) and against the Day of the Month you have the Hour of the Night.—Or, you may find the Hour of the Night by only stepping out, and observing what Star is on or near the Meridian, either above or below the Pole: Then set the same Star on the inner Circle to the same Position, and the Day of the Month will point the Hour of the Night as before.

By this Instrument you may tell, at any Time, by Day or by Night, the Situation of those Stars when they are not visible, by Reason of the Sun's superior Light, or Clouds: For, having set the Day of the Month to the Hour, by your Watch, hold up the Book with its Back towards the North, (as in the last Example) and then the Stars upon the Clock will correspond with those in the Heavens, and point their Places there.

Thus, on the 23d of August, at 1 o'Clock in the Asternoon; I find the Great Bear South above the Pole; the two bright Stars in the Middle of his Back, call'd the Pointers, are upon the upper Part of the Meridian; the other two bright Stars, with the three in his Tail, lie a little to the East.—The Little Bear, and the Dragon's Head full East.—Cepheus and Cassiopeia, near the Meridian, below the Pole.—And Auriga, nearly West.

This Clock will answer, at all Times, these Problems on this Side the Equator: But for those Places beyond it, in South Latitude, the Constellations round the South Pole should be inserted; and in making the Observations, the Book should be held up, and turn'd towards the South Pole.

COELESTIAL GLOBE AND SPHERE.

THE

DESCRIPTION AND USE

OFTHE

COELESTIAL GLOBE.

Stars are placed at proportionable Distances from each other, just as they are in the Heavens; together with the principal Circles of the Sphere.—The Use of this curious Instrument is to point out the various Motions and Appearances of the Sun, Moon, Planets, and Stars, in an easy and natural Manner, without having Recourse to long and tedious Calculations.

The better to determine the Places of the Heavenly Bodies, the several Circles which are describ'd on the Terrestrial Globe, are imagin'd to be transferr'd to the Heavens and delineated among the Stars, exactly over those Circles conceiv'd to be drawn on the Earth.———Thus, opposite the two Poles of the Earth are the two Poles in the Heavens; being the two places where the Earth's Axis (if continued thither) would feem to terminate among the Stars.—Over the Equator on the Earth runs the Equinoctial in the Heavens: —At 232 Degrees on the North Side. runs the Tropic of Cancer, and at 23- South runs the Tropic of Capricorn.—At the same Distance from the Poles are the two Polar Circles, exactly over those on the Earth.—Besides these Circles, there is an oblique one going from one Tropic to another, cutting the Equinoctial in two opposite Places. This Circle is call'd the ECLIPTIC; and is that Path which the Earth would seem to describe among the Stars, in one Revolution, if it were beheld from the Sun. It is divided into 12 equal Parts, and each of these into 30 more, corresponding to the 12 Months of the Year; which 12 Parts are sometimes call'd the 12 Signs, from the Signs or Pictures drawn upon the Stars there *.

Their Names and Characters are here express'd.

Aries v, Taurus v, Gemini n, Cancer n, Leo si, Virgo m,

Libra n, Scorpio m, Sagittarius 1, Capricorn v, Aquarius m, Pisces x.

The first six are call'd-Northern Signs, as lying on the North Side of the Equinoctial and the last six—Southern Signs, because they lye on the South Side the Equinoctial—On this Circle are found the Sun's, the Moon's, and the Planet's Places; that is, under what Degree of any of these 12 Signs or Constellations they appear to be—any Day at Noon.

^{*} Sometimes these Circles are cut out, and put together in form of Rings, intersecting each other as they do in the Heavens, with a little Earth on the Axis in the Center: This serves to give a better Idea of the Position of each Circle than a long verbal Description would do. Their Names and Situation are express'd in the Figure under the Globe.

The Globe, thus delineated *, is affix'd by its Poles to a Brass Meridian, graduated into four Quadrants, as that of the Terrestrial Globe is; (i. e.) on the upper Part—from the Equinoctial to the Poles; but on the under Part from the Poles to the Equinoctial; and is then set in a Wood Horizon, on which are inscrib'd the 12 Months, and the 12 Signs of the Zodiac, the several Degrees of which beingplac'd against the Days of the Month on which the Sun is found in that Part of the Heavens; and the 32 Points of the Compass.—To compleat the whole, there is added a Quadrant of Altitude, to be screw'd on occasionally to the Top of the Meridian. Also an Hour Circle round the North Pole, and an Index.

The Globe being thus fitted up, and the Learner having acquainted himself with the Position and Division of the several Circles and Constellations upon it, may proceed, with Ease and Pleasure, to the Solution of the following

PROBLEMS.

PROBLEM I. To find the Sun's Place in the Ecliptic:

Look the Day of the Month on the Horizon, and against it you have the Sign and Degree of the Ecliptic the Sun is in that Day: Then look the same Sign and Degree on the Ecliptic Line on the Globe, and there fix on a little black Patch, which will represent the Sun's true Place, and shew what Stars are round him that Day.

Thus, on the 24th of May the Sun is in $3\frac{1}{2}$ Degrees of Gemini, or the Twins:
—Is situate near the Bull's Eye and the Seven Stars, which are not then visible by Reason of his superior Light. But were the Sun to suffer a total Eclipse on that Day, you would then see those Stars shining near him very bright.

PROBLEM II. To find the Sun's Declination.

Having found the Sun's Place, (as in the last Problem) bring it to the Bress Meridian; and the Degree which stands over it, shews the Sun's Declination, or Distance from the Equinoctial, either Northward or Southward at that Time.

Thus, on the 24th of May, you will find the Sun to have 21 Degrees of North Declination: But on the 12 of December, he is found to have $23\frac{1}{4}$ Degrees of South Declination.

* The convex Surface of the Globe corresponds with the concave Surface of the Heavens; whence it is easy to conceive, that, if Holes were made in each Star, and an Eye plac'd in the Center of the Globe, it would view, thro' those Holes, the same Stars in the Heavens represented by them.

The Heavens are studded over with Stars as much by Day as by Night, only the greater Light of the Sun renders them invisible to us. But when that Luminary descends below the Horizon, they begin gradually to appear. When the Sun is gone down about 12 Degrees, Stars of the first Magnitude become visible. When he is 13 Degrees below the Horizon, those of the fecond Magnitude are seen. When 14 Degrees, Stars of the third Magnitude appear. When 15 Degrees, those of the fourth Magnitude present themselves to view. When he is descended 16, 17, and 18 Degrees, Stars of the fifth and sixth Magnitude, with those of the smallest Size, become conspicuous, and the azure Arch sparkles with all its Glovy.

PROBLEM III. To rectify the Globe, or to adjust it so to the Place where you are, as it may represent the present State or Situation of the Heavens.

Raise the Pole, till it stands as many Degrees above the Horizon as is the Latitude of the Place you are in; which may be done by the Degrees mark'd on the lower Side of the Meridian, and number'd from the Pole: Fix the Quadrant of Altitude on the Meridian, to the Zenith; that is, fix it on the Top to the Latitude of the Place: Turn the Globe, by the Help of a Compass, till the Pole points toward the real Pole of the World: Bring the Sun's Place in the Ecliptic to the Meridian, and set the Index to the upper 12; so will the Globe exhibit the Face of the Heavens for the Noon of that Day, at the Place you rectified the Globe for: And by turning the Globe till the Index points to any given Hour, you will see the Situation of the Heavens there at that Hour.

If the Place be in North Latitude, the North Pole must be rais'd above the Horizon; if in South Latitude, the South Pole must be rais'd.

PROBLEM IV. To find the Time of the Sun's Rising and Setting; and also his Amplitude, or Distance from the East or West Points of the Horizon, at the Times of his Rising and Setting; and the Length of the Day and Night.

The Pole being elevated to the Latitude of the Place where you are, bring the Sun's Place in the Ecliptic to the Brass Meridian, and set the Index at the Pole to the upper 12 or Noon: Then bring down the Sun's Place to the Eastern Side of the Horizon, and the Number of Degrees between the East Point and the Point where the Sun is,—is the Amplitude at Rising. The Hour pointed to by the Index on the Dial-plate, is the Time of the Sun's Rising. The Sun's Place brought to the Western Edge of the Horizon, shows there, in like Manner the Amplitude at Setting; and the Index points out the Time of the Sun's Setting.

The Time of the Sun's Rising being doubled gives the Length of the Night:

And the Time of his Setting being doubled gives the Length of the Day,
or the Time he appears above our Horizon.

Thus, on the 24th of May, at London, the Sun rises at 4, with 36 Degrees of Eastern Amplitude (i. e. 36 Degrees from the East to the North); and sets at 8, with 36 Degrees of Western Amplitude; making the Day 16 Hours long, and the Night 8 Hours.——The Sun now rises upon the North-East by East Point of the Combass, and sets upon the North-West by West Point.

PROBLEM V. To find the Beginning and End of Twilight.

Restify the Globe; turn it Westward, with the Quadrant of Altitude, (which must be screw'd on the Meridian to the Latitude of the Place) till the Point opposite the Sun's Place in the Ecliptic cuts the Quadrant at 18 Degrees above the Horizon; the Index will then shew the Time Twilight begins in the Morning.—The tame Point, brought with the Quadrant of Altitude to the Eastern Side the Globe, till it cuts the 18th Degree, as before, the Index will then point the Time when Twilight ends in the Evening*.

Thus, on the 24th of May, you will find the Point in the Ecliptic opposite the Sun's Place, (which is $3\frac{1}{2}$ Degrees of Sagittarius) will not reach the 18th Degree in the Quadrant of Altitude, which shows that the Sun does not descend, at that Time, 18 Degrees below the Horizon; consequently, we have then no Night, but a continued Twilight.—But, on the 12th of December, Twilight begins about Half past 5, and ends about Half past 6. And on the 21st of March, and 23d of September, Twilight begins about 4 in the Morning, and ends about 8 at Night.

PROBLEM VI. To find the Azimuth and Altitude of the Sun at any Time of the Day.

The Globe being rectified, and the Quadrant of Altititude screw'd on the Brass Meridian at the Latitude of the Place where you are, bring the Sun's Place to the Meridian, and set the Index to 12: Then turn the Globe, till the Index points the Hour given, and lay the Quadrant to the Sun's Place; so will the Degree of the Quadrant lying upon the Sun, be his Altitude above the Horizon: And the Number of Degrees counted on the Horizon from the South Point to the Point where the End of the Quadrant cuts it, will be his Azimuth at that Time.

Thus, on May 24, at 9 o'Clock in the Morning, you will find the Sun to have about 44 Degrees of Altitude above the Horizon; and $62\frac{1}{2}$ of Azimuth from the South. And at 6 in the Evening of the same Day, his Altitude is 20 Degrees, and Azimuth 79 Degrees, so that the Sun is then upon the West by North Point of the Compass.

^{*} Twilight is that faint Light we have in a Morning before the Sun rises, and in the Evening after the Sun is set. This is found by Observation, to begin when the Sun comes within 18 Degrees of the Horizon in the East, and to end when he is gone 18 Degrees below the Horizon in the West.——In the Southern Parts of England, the Inhabitants have no real Night, but a Twilight, from the 20th of May to the 22d of July.

PROBLEM VII. To find the Oblique Ascension—Right Ascension—and Oblique Descension of the Sun.

The Globe being rectify'd, and the Sun's Place brought to the Eastern Side of the Horizon; the Number of Degrees accounted from the Beginning of Liries to that Degree of the Equator now come to the Horizon, is the Sun's Oblique Ascension. The Sun's Place being brought to the Meridian, the Degree of the Equator lying under it is his Right Ascension. And, by bringing the Sun's Place to the Western Side of the Horizon, the Degree then of the Equator in the Horizon, is his Oblique Descension.

Thus, on the 24th of May, (the Sun being in 3: Degrees of Gemini) his Oblique Ascension will be found to be 32 Degrees—his Right Ascension 62 Degrees—and his Oblique Ascension 91 Degrees.

The Difference in Degrees between the Right and Oblique Ascension turn'd into Time, by allowing 15 Degrees for every Hour, shows how much the Sun (at that Time) rises before or after 6 o'Clock. This is call'd the Ascensional Difference; which Difference is to be subtracted from 6 in the Morning, and added to 6 in the Evening thro' the Summer Half-Year, but added to 6 in the Morning and subtracted from 6 in the Evening thro' the Winter-Half, to give the Time of the Sun's Rising and Setting—In the above Example, the Ascensional Difference is 30 Degrees (which in Time is 2 Hours) to be subtracted from and added to 6; consequently, the Sun rose that Morning at 4, and set that Night at 8.

PROBLEM VIII. To find the Right Ascension and Declination of any Star represented on the Globe.

Bring the Star to the Brass Meridian; observe the Degree which stands over it, for that is the Declination: and the Degree of the Equator which comes to the Meridian with the Star is its Right Ascension.

Thus you will find the Declination of Aldebaran, in Taurus, to be about 16 Degrees North; and the Right Ascension about 65 Degrees.—And, in like Manner, the Declination of Syrius, or the Great Dog, to be about 16; Degrees South; and the Right Ascension about 98 Degrees.

The Right Ascension and Declination of the Sun alter every Day on Account of his Motion in the Ecliptic; but the Right Ascensions and Declinations of the fix'd Stars continue the same throughout the Year.

PROBLEM IX. To find the Longitude and Latitude of a Star or Planet.

Apply the End of the Degrees of the Quadrant of Altitude to that Pole of the Ecliptic which lies on the same Side the Ecliptic with the Star, and bring its graduated Edge to the Star, so will the Degree of the Quadrant at the Star point its Latitude reckon'd from the Ecliptic towards its Pole: And the Degree of the Ecliptic cut by the Quadrant, at that Time, will be the Longitude sought.

Thus, the Latitude of Capella, or the Goat, in the Left Shoulder of Auriga, is about $22\frac{1}{2}$ Degrees North, and Longitude about 19 Degrees in Gemini.——And the Latitude of Fomalbaut, in the Southern Fish, is about 21 Degrees South, and Longitude $29\frac{1}{2}$ Degrees in Aquarius.

- The Latitude of a Star is its Distance in Degrees and Minutes, counted from the Ecliptic Line towards its Pole.—If the Star lie on the North Side the Ecliptic, it has North Latitude; if on the South Side, it has South Latitude.
- The Longitude of a Star is its Distance from (the Equinoctial Point) Aries, reckon'd in Signs, Degrees, and Minutes. If a Star be situate on the North Side the Ecliptic, you must apply the Quadrant to the North Pole of the Ecliptic, but if on the South Side you must apply the Quadrant to the South Pole of the Ecliptic, as was done in the above Example of Fomalbaut.

PROBLEM X. To find the Oblique Ascension and Oblique Descension of any Star.

Elevate the Globe to the Latitude of the Place where you are; then bring the Star to the Eastern Side of the Horizon; and the Degree of the Equator then cut by the Horizon is the Oblique Ascension: Turn the Star to the Western Side, and the Degree of the Equator cut by the Horizon is the Star's Oblique Descension.

Thus at London you will find the Oblique Ascension of Aldebaran, or the Bull's Eye, to be about $43\frac{1}{2}$ Degrees; and the Oblique Descension about 87 Degrees.—In like Manner, the Oblique Ascension of Syrius, or the Great Dog, will be found to be about 121 Degrees, and the Oblique Ascension about 75 Degrees.

The Sun's and a Planet's Oblique Ascension and Descension differ every Day in the same Latitude by Reason of their Motion in the Zodiac, but a six'd Star's Oblique Ascension and Descension are every Day the same thro' the Year.

PROBLEM XI. To find the Rising, Southing, and Setting of any Star, together with its Continuance above the Horizon.

Rectify the Globe for the Place where you are; bring the Sun's Place to the Meridian, and let the Index to the upper 12: This being done, bring the Star you enquire after to the Eastern Side of the Horizon; and the Index will shew the Hour of its Rising.—Turn it on to the Meridian, you have (by the Index) the Time of its Southing.—And, by bringing it to the Western Side of the Horizon, the Time of its Setting.—The Hours on the Dial-Plate from the Rising to Setting, show the Continuance of the Star above the Horizon, and the remaining Hours, the Time the Star is below, or under the Horizon.

Thus, on May 24, at London, the Star Aldebaran, or the Bull's Eye, will be found to rise about three Quarters past 4 in the Morning; will come to the Meridian about Noon; and set about 7 in the Evening. It therefore continues above the Horizon 15 Hours, and below it 9 Hours.

This Star is not visible at that Season of the Year, by Reason of its Near-ness to the Sun.—But, on the 24th of November, at London, Aldebaran will rise about three Quarters past 4 in the Evening, and set about three Quarters past 7 in the Morning; so that it will be then visible all Night, if the Sky continue clear.

PROBLEM XII. To find the Eastern and Western Amplitude of a Star.

The Globe being rectify'd, as before, and the Star brought to the Eastern Side of the Horizon, you have, upon the Horizon, the Amplitude, i. e. the Distance in Degrees from the East, either Northward or Southward; and also, the Point of the Compass the Star rises upon. Turn the Star to the Western Side, and the Horizon shews the Amplitude and Point of the Compass it sets upon there.

Thus, on the 24th of May, at London, Aldebaran will have about $25\frac{1}{2}$ Degrees of Amplitude from the East towards the North at Rising; and about $25\frac{1}{2}$ Degrees of Amplitude from the West towards the North at Setting——It rises nearly East-North-East, and sets nearly West-North-West.

The Stars, though they rise at different Hours of the Day, yet always rise on the same Points of the Compass.

PROBLEM XIII. To find the Altitude and Azimuth of a Star at any Hour.

The Globe being rectify'd, and the Quadrant of Altitude screw'd on the Meridian at the Latitude where you are: Bring the Sun's Place to the Meridian, set the Index to 12, and turn the Globe till the Index Points the Hour given: Then by laying the Edge of the Quadrant upon the Star, you find (by the Degree against it) the Altitude above the Horizon at that Time; and the End of the Quadrant will shew, upon the Horizon, the Azimuth of the Star, and also, the Point of the Compass it then bears upon from you.

Thus, at London, on November 24, at 9 at Night, you will find, (by proceeding as above) that Aldebaran will have about 36; Degrees of Altitude; with 68 Degrees of Azimuth from the South; and that it bears upon the East South-East Point of the Compass.—But at 3 in the Morning, he will have about 41 Degrees of Altitude; about 59 Degrees of Azimuth; and bear upon the South-West by West Point of the Compass, nearly.

PROBLEM XIV. The Altitude of the Sun by Day, or of any Star by Night being given, to find the Hour of the Day or Night.

Rectify the Globe to the Place where you are; fix on the Quadrant of Altitude to the Meridian at the Latitude given; bring the Sun's Place, at that Time, to the Meridian, and fet the Index to 12. This done, turn the Globe and the Quadrant together, till the Star, or the Degree of the Ecliptic the Sun is then in, cuts the Quadrant in the Altitude given; the Index will then point the Hour fought.

Thus, on May 24, at London, when the Sun is observed to be about 44 Degrees bigh, in the Morning, you will find the Time to be 9 o'Clock. And when his Altitude is 20 Degrees, in the Afternoon, you will find it to be 6 at Night.—Also, on November 24th, at Night, Aldebaran being observed to be about 36½ Degrees bigh, and to the East of the Meridian, we find (by proceeding as before) that it was 9 o'Clock when that observation was made.

PROBLEM XV. To measure the Distance, in Degrees, between any two Stars.

Lay the Beginning of the Quadrant of Altitude upon one of the Stars; apply the graduated Edge to the other; and the Degrees contain'd between the Stars are their Distance requir'd.

Thus the Distance between Aldebaran and Capella will be found to be 30 Degrees. And the Distance from Syrius (the Great Dog) to Deneb (the Lion's Tail) 81 Degrees.

The same might have been done with a Pair of Dividers; for opening from one Star to another, and applying the Wideness to the Equator, the Degrees between the Legs is the Distance, as before. If the Quadrant is not long enough to reach from one Star to the other at once, you may measure the Distance at twice.

PROBLEM XVI. To find what Stars never rife nor set to the Inhabitants of any given Place.

All those Stars toward the North Pole, which lie not further from it than the Number of Degrees equal to the Latitude of the Place you are in, never set at that Place.——And all those Stars which lie within the same Number of Degrees from the South Pole never rise, or appear above the Horizon of that Place.

Thus, at London, if you elevate the Pole to the Latitude of 51½ Degrees, you will see, by turning the Globe round, that all the Stars which are not further from the North Pole than 51½ Degrees, as the Great Bear, Little Bear, Draco, Cepheus, and Cassiopeia, never set, or descend below the Horizon:——And also, that those Stars situate at the same Distance from the South tole, as Hydrus, Toucan, Indus, Pavo, South Triangle, Musca, Crossers, Pisces volans, &c. never ascend above the Horizon of London, but remain ever invisible to the Inhabitants there.

Hence it is easy to conceive, that the Inhabitants at the North Pole can never see any of the Stars lying South of the Equinoctial; nor can the Inhabitants at the South Pole see any Stars lying North of the Equinoctial. But those who inhabit at the Equinoctial will have the Pleasure of seeing all the Stars from Pole to Pole.

* If a Circle be drawn round the Stars which never set; and another round those which never rise in the Latitude you are in; the former is call'd the Circle of perpetual Apparition; and the latter, the Circle of perpetual Occultation: All the Stars situate between these two Circles rise and set every Day.

PROBLEM XVII. To distinguish the Stars in the Heavens one from another, and to know them by their proper Names.

Rectify the Globe to the Latitude of the Place where you are; bring the Sun's Place to the Meridian, and fet the Index to 12: Then turn the Globe to the given Hour, and there let it remain; so will every Star inscrib'd upon it exactly correspond with, and point to the same Star in the Heavens; and by transferring the Eye from the Globe to the Stars, you will evidently discover the same there *.

Thus, on the 24th of November, at 10 at Night, (the Globe being rectify'd, and plac'd, by means of a Compals, North and South) you will immediately perceive the North Pole of the Globe point to a Star of the second Magnitude, which is the Pole-star; round which all the rest circulate every 24 Hours.—At the same Time you may observe, a little lower, two Stars, not quite so bright as the Pole-star, almost in a Right-Line with it, and sour more which form a Kind of Quadrangle: These seven make the Constellation call'd the Little Bear, the Pole-star being in the End of the Tail ——, Near the same Place, but more towards the East, you may observe seven bright Stars, all of the second Magnitude, which are commonly call'd Charles's Wain but by Astronomers the Great Bear; the two foremost of the Square lie almost in a Right line with the Pole star, and are called the Guards, or Pointers; so that knowing the Pointers, you may easily find the Pole-star.——

^{*} If a Right Line be imagin'd to proceed from the Center of the Globe, (when rectify'd) through any Star on its Surface, that Line will fall upon the same Star in the Heavens, and point it out, if continued thither

Further on, almost East, and near the Horizon, is seen a bright Star of the first Magnitude, call'd Procyon, or the Little Dog. ————Higher up the Sky, and a little more to the North, you find two bright Stars in Gemini, or the Twins; the uppermost is of the first Magnitude, call'd Caster, the other of the second Magni-ant Star Capella, or the Goat, in the Shoulder of Auriga; this is of the first Magnitude.——Lower down, and nearly South-East, you observe Aldebaran, or the Bull's Eye, a Star of the first Magnitude, and near that the Pleiades, or noted Seven Stars (all small ones) in the Back of the Bull.———Between Taurus, or the Bull, and the Horizon, is seen the glorious Confellation Orion, with two bright Stars of the first Magnitude; one in his Right Shoulder, call'd Bed elgueze, the other in his Left Heel, call'd Regel: In the Middle of this Constellation are three Stars of the second Magnitude in a Right Line, and a little lower, three others of the third Magnitude, lying also in a strait Line with each other; which Stars, by the Country People, are call'd the Ell and the Yard, but by Astronomers, Orion's Belt and Sword.

In this Manner the rest of the Stars, in any other Constellation may be readily found, by observing how they lie with respect to those already known. A little Practice will make you well acquainted with them all.

Between the Bull's Eye and the Seven Stars, you observe a large Star of a heavy dull Colour; but upon the Globe you don't find any, in that Place, corresponding with it. This shews that Star to be one of the Planets; and, by looking into Mr. White's Ephemeris, you will perceive it must be Saturn: For, Saturn, at that Time, is in about 2 Degrees of Gemini, with about 2 Degrees of South Latitude, which Situation exactly agrees with that Part of the Heavens where you see him.

PROBLEM XVIII. To find all those Stars, which at any Hour, are Rising, Setting, or upon the Meridian.

Having rectify'd the Globe, as in the last Problem, turn it till the Index points to the given Hour. The Globe being kept in that Position, all those Stars in the Eastern Edge of the Horizon are then Rising, those under the Brass Meridian are upon the Meridian in the Heavens; and those on the Western Edge of the Horizon are setting.

Thus, on the 24th of November, at 10 at Night, we find Charles's Heart, the Lion's Head, and Syrius, or the Great Dog, rising.—The Little Bear, Perseus, the Ram and Whale, upon or near the Meridian.—And Aquarius, or Water-Bearer, the Eagle, Hercules, &c. setting.

PROBLEM XIX. To find the *Place* of any Planet upon the Globe, and by that Means, its true *Situation* in the Heavens.

First, by some Ephemeris (Mr. White's is one of the best) which has the Planets Places calculated to every Day at Noon, find the Longitude and Latitude of that Planet: then make a Mark with a Pencil, and stick a little black Patch in that Point on the Globe, which Point will shew its true Situation, with respect to the Stars, and also what Stars immediately accompany it.

Thus, November 24, 1765, I find, by Mr. White's Ephemeris, that Saturn is in about 2 Degrees of Gemini, with a little more than 2 Degrees of South Latitude, at which Place I fix on a little black Patch, representing his Situation in the Heavens.—The same Day I find Jupiter to be in 22½ Degrees of Leo, with three Quarters of a Degree of North Latitude, where I stick on another Patch for his true Place. In like Manner you may set on little Patches representing all the Planets, which will shew you their true Places at that Time, and how they are situate with respect to the Stars, and to one another.

PROBLEM XX. To find when any of the Planets rise or set, or come to the Meridian; as also, the Point of the Compass they rise or set upon.

Rectify the Globe to the Latitude where you are; bring the Sun's Place to the Meridian, and fet the Index to the upper 12, and the Globe will exhibit the exact Situation of the Heavens at Noon that Day. Then bring the little Mark representing the Planet to the Eastern Side of the Horizon, and the Index will point the Hour of its Rising: turning it on to the Meridian, you will see the Time of its Southing; and bringing it down to the Western Side, you see the Index point the Hour of its Setting.

Thus, November 24, 1765, Saturn rifes nearly North-East by East, about a Quarter past 4 in the Evening; comes to the Meridian, or souths about Midnight; and sets nearly North-West by West, about half past 7 next Morning. Jupiter rises East-North-East, about a Quarter past 10 at Night; comes to the Meridian at Half past 5 next Morning; and sets West-North-West at 1 in the Asternoon. After this Manner may the rising, setting, &c. of the rest of the Planets be found.

The Azimuth, Altitude, and Declination of the Planets, and also their Amplitudes at rising and setting, are found in the same Manner as those of the Sun and Stars delineated in the foregoing Problems.

PROBLEM

PROBLEM XXI. To find the Cause of a Comet, and also the Position and Direction of its Tail in the Heavens; with its Rising, Setting, Amplitude, Azimuth, &c. at the Time whilst it continues in our Part of the System.

When a Comet is feen in our Hemisphere, observe its exact Situation with respect to the Stars about it, and likewise the Form and Length of its Tail, and stick a small Piece of Paper upon that Part of the Globe, cut into like Form with the Comet. Do so for two or three Nights successively, and set on like Patches, representing his Place at each Observation. By this Means you will discover its Course, which will be always in a great Circle nearly: for a Thread laid over these Patches round the Globe, will exhibit the tract it takes in the Heavens, and the Constellations it passes through.—Where the Thread cuts the Ecliptic Line will be the Place of its Nodes.

Now if the Globe be rectified for the Day you make the Observation, the Sun's Place brought to the Meridian, and the Index set to 12, you may find the Time of the Comet's rising or setting, by bringing it to the Eastern and Western Side of the Horizon. You may also find its Azimuth, Amplitude, Altitude, Longitude, Latitude, &c. (Position of its Tail with respect to the Horizon any Hour) in the same Manner as directed before for the Stars.

TOMAKE

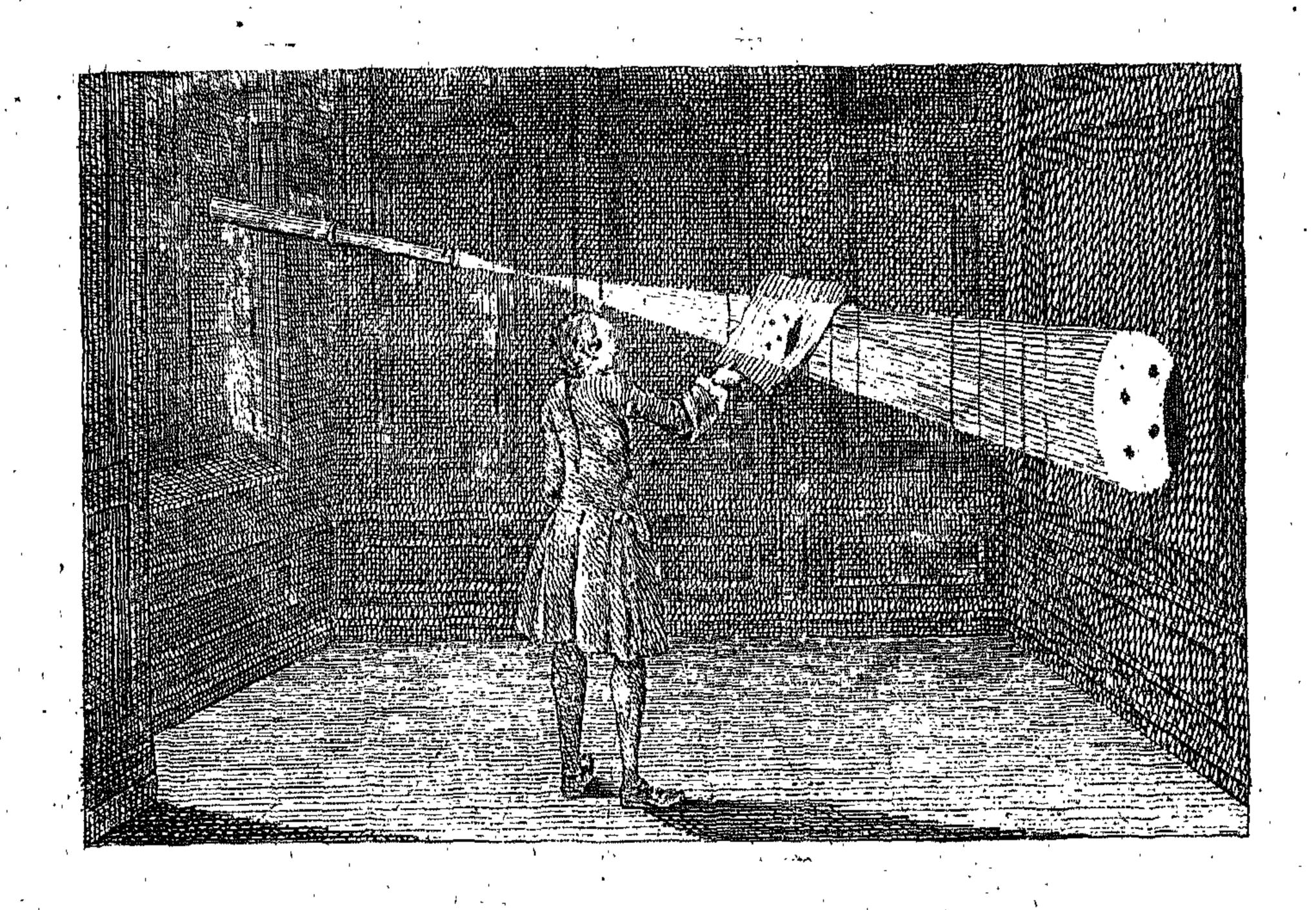
CURIOUS OBSERVATIONS

UPON THE

SUN, MOON, AND ECLIPSES,

IN A DARKENED ROOM.

THE Room you make choice of for Experiments of this Kind, must have a Window in it, which opens towards the Object you would observe: let it also be made quite dark, by means of a Shutter placed before the Window; then assix the Object-End of the Telescope of about 4 or 5 Feet in Length into a Hole made in the Shutter; draw out the other End to its usual Length, and turn it about till the Sun's Rays, passing through the Glass, fall upon a Sheet of white Paper held a little beyond the Focus, and you will have represented on the Paper the exact Image of the Sun, as it really is in the Heavens, with the Spots distinct upon his Face at that Time, as appears in the Figure below.



If the Paper be removed further from the Glass, the Image will increase its Dimensions: and it may be so contrived, as to fall on the opposite Wall (if it be covered with any Thing white), where the Image will be magnified several Feet, and the Spots an Inch, or more, in Diameter.

The Moon, likewise, may be exhibited in this Manner, but her Light

is rather too faint to render the Spots very distinct.

But by this Method an Eclipse of the Sun is seen to the greatest Advantage; for without any Hurt or Uneasiness to the Eye, you behold the whole Phænomenon of the Eclipse; how it begins, and where it ends; in what Manner the Moon passes over the Sun's Disk, with the Quantity eclipsed at every Moment, which is a most agreeable and delightful Appearance.

ASTRONOMICAL

ASTRONOMICAL PARADOXES.

AVING at the End of our View of the Earth inferted a few Propositions (which are generally called Paradoxes, from the Surprise or Fallacy they seem to carry with them at first Sight), purely to call the Attention of the young Geographer to a Reslection on what he has been perusing; we have here added a few more for the Exercise of the young Astronomer; the Solution of which may be deduced from the Nature and Consideration of the System, delivered in the foregoing Pages.

PARADOX I. There are feveral Places upon the Globe of Earth, where the Sun and Moon, and all the Planets, do actually rife and fet according to their various Motions, but never any of the fixed Stars.

SOLUTION. Under the Poles, the Planets, by their Revolutions in their Orbits, are sometimes North, sometimes South of the Equinoctial Line; consequently rise and set as the Sun does, with respect to these Places. But the fixed Stars keeping an exact Distance from the Poles, may be said never to rise or set, but continually move round them.

The Motion which the fixed Stars have upon the Poles of the Ecliptic, being so very slow as to make no visible Alteration in their Places for Ages, cannot be any Objection to this Paradox.

PARADOX II. There is a very remarkable Place upon our Globe, where all the Planets, notwithstanding their different Motions and various aspects, do all bear upon the same Point of the Compass.

SOLUTION. Under either of the Poles.—For to an Eye at the North-Pole, all the Stars, however fituate, will bear upon the South Point: and, at the South Pole, they will, on the contrary, bear on the North Point of the Compass; because every Point of the Compass becomes a Meridian, as observed in Paradox I. in my Geography.

For the same Reason, and at the same Places, a full Moon, which is always opposite the Sun, may both happen to rise and set at the same Instant of Time, and upon the same Point of the Compass.

PARADOX III. There is a certain Place in the Island of Great Britain where the Stars are always visible, at any Time of the Day, if the Sky be not overcast with Clouds.

SOLUTION. The Place may be some deep Pit or Well, or Professor Maskeline's Observatory, at Greenwich, such as Tycho Brache had in Denmark, which was a deep Well beset with Glasses, where he sat and observed the Stars at all Seasons.

PARADOX IV. There is a certain Island near the Continent of Europe, fome of whose Inhabitants are of such exquisite Sight, that even with one of their Eyes, they can actually behold 10 Moons, real and true, all at once above the Horizon; and also, Ten Times the Number of Stars beheld by others with both their Eyes at the same Instant.

SOLUTION. Any Island, or other Place, provided the Observer is surnished with a good Telescope, which will shew the 5 Moons of Saturn, the 4 of Jupiter, and our own Moon, all which are 10 in Number. Also they will see through their Glass 10 Stars for 1 beheld by the naked Eye. Nay, in the Milky Way, where we cannot behold even one Star, they will see an infinite Number so close and thick together, that them united, though a faint Light, causes that whitishness which gives it the Denomination of the Milky Way before-mentioned.

PARADOX V. To several Parts of the Globe there are several Planets, which are so far from coming to an Opposition, that they form neither Square, Trine, nor Sextile Aspect with the Sun.

SOLUTION. Venus and Mercury, whose Orbits are contained within the Orbit of our Earth, never form any of these Aspects with the Sun.

PARADOX VI. There are feveral *Planets*, faid to be in Conjunction with the Sun, not only when they appear to be in the same Degree of their Orbit with the Sun, but when they are in that Degree of their Orbit diametrically opposite the former.

SOLUTION. The several Planets are Venus and Mercury, who have a twofold Conjunction with the Sun, one in the superior and another in the R

inferior or opposite Point of their Orbits. In the inferior Conjunction they pass like Spots between the Sun and us; but in the superior Conjunction, they pass behind the Sun, leaving him between us and them.

PARADOX VII. There is one certain Place in the Universe, where the Planets, both inferior and superior, may be constantly seen to move forwards in the same regular and uniform Manner, though, to most Places of the Earth, they appear, at the same time, to move very unequally; and sometimes they seem to run back, and at other times not to move at all.

SOLUTION. At the Sun, the Center of the System, the Planets move all regular and direct according to the Order of the Signs; but to us, out of the Center of their Orbits, they appear sometimes direct, sometimes stationary, and sometimes retrograde.

PARADOX VIII. Though the Number of the Stars in a clear Winter's Night seems almost infinite; and Mr. Flamstead (late Astronomer Royal) has given us a Catalogue of 3000, yet it is difficult for the most penetrating Eye to reckon at any one Time above 1000

SOLUTION. This Appearance is only a Deception of the Sight, arising from their vehement twinkling, whilst we look on them confusedly, and without reducing them to Order. For, when we come to view them more distinctly, we find the confused Number vanish, and the whole Number of Stars discoverable by the naked Eye is not above a Thousand.

Though the Number of Stars which can be seen by the naked Eye are so few, yet it is probable there are many more beyond the Reach of our Sight. For, through good Telescopes, they appear in vast Multitudes in every Part of the Heavens, and the better the Glasses are, the more are still discovered. The learned Dr. Hook has observed 78 Stars in the Pleiades, of which the naked Eye cannot discern more than 6. And in Orion, which has but 80 Stars in the British Catalogue, there has been counted more than 2000.

PARADOX IX. There is a certain Place upon the Earth, above whose Horizon Saturn is 15 Years together; and there is another Place of considerable Distance from the former, which has Jupiter nearly 6 Years above their Horizon, without once setting during that Time.

SOLUTION. This Place is under, or near, the Poles, where Saturn continues without setting all the Time he is in that Half of his Orbit, North of the Equinoctial, which is about 15 Years, his Period being 30 Years. The other Place at a considerable Distance is the South Pole, where Jupiter must continue visible through that Part of his Orbit South of the Equinoctial, which is about 6 Years, his Period being 12 Years nearly.

PARADOX X. There are feveral Planets, or wandering Stars, which at certain Times appear and disappear, whose light decreases as they come towards the Earth, and increases as they go from the Earth; and so transparent are they, that the smallest Stars can be seen through them. There are also others so opaque, that in Conjunction with the Sun, they appear as Spots on his Face; and the farther they are from us, the bigger they seem.

SOLUTION. The first Part of this Paradox respects the Comets, which wander through all the Orbits of the Planets, and to vast Distances beyond Saturn, where they become invisible. These wonderful Bodies decrease in their Light as they come towards the Earth (if they pass between us and the Sun), and increase as they go from it, by turning more of their illuminated Side towards us. Through their amazing Tails, which are so very rare and thin, Stars of the smallest Magnitude may be seen. The other opaque Planets, which appear as Spots on the Sun's Face, are Venus and Mercury in their inferior Conjunctions, passing between us and the Sun; and the Moon, every one knows, appears much bigger in the Horizon than in the Meridian, though, in the latter Situation, she is above 4000 Miles nearer us than in the former.

This Deception is owing to the Refraction of the Atmosphere, as observed in one of the Paradoxes of my Geography.

To these Paradoxes Numbers might be added, but they would swell the Book beyond the Dimensions of a Compendium, as well as, perhaps, tire the Youth in the Perusal. I shall, therefore, only beg leave to detain him, whilst I offer to his Attention an Observation on the beautiful and pious Exclamation of royal David on viewing the Heavens.

" When I behold (says he) the Moon and the Stars that thou hast ordained, full of various Beings, whose Natures and Constitutions are admirably adapted to the World thou hast placed them in; who have still, perhaps, retained their original Innocence and Purity; whose Hearts are now warm with Reflections on thy Goodness, and whose Tongues are telling forth thy Praise: Lord what is Man! Man who has so often broken thy Laws, and ungratefully disregarded thy holy Statutes, that thou art mindful of him! that thou dost continue him still in Existence, and furnish him so liberally with every blessing of Life! or what is the Son of Man, the poor Inhabitant of this small Corner of thy infinite Creation, that thou so regardest him! that thou hast in fuch a singular Manner procured him a Redeemer, who has so happily de-· livered him from the Miseries justly due to his Ingratitude, and conferred on him Immortality and Glory for ever."

 $*_**$ The Learner having carefully gone over this System, will read Dr. Keil's Lectures, Dr. Young's Astronomy, and Mr. Martin's Young Lady's and Gentleman's Philosophy with a peculiar Pleasure.